Characteristic Features of the Scattering of Elementary Particles Arising from their Spin and Electromagnetic Interaction

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The work presented here is claimed as original except where explicit reference is made to the work of others and no part has been submitted for any other degree or diploma.

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CONTENTS

		Page
CHAPTER 1	Introduction	1
CHAPTER 2	The Analyticity of Scattering Amplitudes	
	for Particles with Spin	4
2.1	General Principles	4
2.2	Derivation of the Trueman-Wick Crossing	
	Relation	7
2.3	Derivation of Invariant Amplitudes	15
2.4	Comparison of the Various Amplitudes	
	used for Spinning Particles	21
	Appendix	30
CHAPTER 3	Partial Wave Amplitudes, Pole Residues	
	and Form Factors	37
3.1	Introduction	37
3.2	Partial Wave Amplitudes	38
3.3	Pole Residues	48
3.4	Form Factors	. 53
CHAPTER 4	Regge Theory of Resonance Production	60
4.1	Introduction	60
4.2	Regge Theory in Leading Order	61

Introduction

Nost theoretical treatments of the dynamics of strongly interecting particles confine themselves to spinless particles and assume that the introduction of spin into the problem brings only incessential complications. Indeed in most cases spin sets simply as an extra label and may be taken into account by a multichansel formulation. However one must first choose which of the many possible amplitudes, differing by transformations on the spin indices, one will use. As shown in many papers, an unmise choice of amplitude can lend to long calculations and makes a simple generalization appear very complicated. This is particularly so when invariant amplitudes are used without considering whether the information thereby expressed is relevant to the problem at hand.

Recently libra and finag have suggested a new way of studying the snallyticity of scattering applitudes for particles with spin. The nost important feature of their method is the observation that kinematic singularities arise from the singularities of the lowests transformations generating the chosen states. By relating our chosen amplitudes to those generated by Lorents transformations without these singularities, one can expose the kinematic singularities. The Trumsmitck crossing relation is a particular case of such a transformation, It is characteristic of this method that it only exposes the singularities locally, and not everywhere as can be done through a suitable choice of invariant amplitudes. However this is usually sufficient as will be seen from the examples considered in this themis.

In Chapter 2 we give the theory of this method and compare it to the approach via invariant amplitudes. This will concentrate on the full empittude, while in Chapter 3 we apply the theory to the residuce of (fixed) poles in scattering molitudes, partial wave molitudes and form factors.

This is extended in Chapter 4 to a treatment of Regge poles. This necessitates a discussion of daughters and conspirators, but they are only considered in enough detail to enaulte one to apply the theory to the experimental situation, Preliminary results on some fits to the production of $\frac{1}{2}$, $\frac{1}{2}$ (136), $\frac{1}{2}$ (136) at medium and high energy are presented.

Throughout this work we shall make the nontrivial restriction to non-zero mass particles and in the Regge section

Finally in Chapter 5, which has no connection with the previous work, we consider two practical problems in combining electromagnetism and strong interactions. These are Dashen's calculation of the proton-neutron mass difference and the extraction of Coulomb-nuclear interference from experimental data on elastic scattering at high energies.

The Analyticity of Scattering Amplitudes for Particles with Spin

2.1 General Principles

In this chapter, the theoretical background to the work of chapters 3 and 4 will be described.

First in Section 2.2, we will give a derivation of the Trueman-Mick (hereafter called TM) crossing relation which is sufficiently complete to obtain the unknown phase in their formula. To make this worthmalle, requires a rather careful definition of the states and amplitudes to be used. This has been given by \$tap\$^{\$1}\$ and Taylor\$^{72}\$ and I base my notation on their work. In the appendix to this chapter I have collected all definitions and conventions used in this thesis. Here, for instance, M functions and helicity amplitudes are defined.

The study of the analyticity of scattering amplitudes for particles with spin may be based on the principle

(P) M functions are analytic functions of momenta, and kinematic

singularities occur through the choice of their arguments $p_{\underline{i}}=p_{\underline{i}}(s,t) \quad \underline{i}=1 \, \ldots \, 4 \quad \text{to obtain a function of s and t.}$

From the relation (2.A.7), given in the appendix, between helicity amplitudes and M functions, we may derive S^{2} the alternative form of (P) ...

(P') The helicity amplitudes $H_{u}^{\lambda_{2}\lambda_{3}:\lambda_{2}\lambda_{1}}$ (x,y) {e.g. x=s, y=t}, where x is energy and y is momentum transfer, are for fixed x, analytic functions of y, except at the physical region boundary where they behave like

$$\mathbb{E}_{\mathbf{x}}^{\mathbf{\lambda}_{4}\lambda_{5}:\lambda_{2}\lambda_{1}} \sim \left[\frac{1-\cos\varphi_{\mathbf{x}}}{2}\right]^{\frac{1}{2}\left|\lambda_{1}-\lambda_{f}\right|} \left[\frac{1+\cos\varphi_{\mathbf{x}}}{2}\right]^{\frac{1}{2}\left|\lambda_{1}+\lambda_{f}\right|}$$
(2.1.1)

This latter behaviour may for instance be derived from (2.A.7), and the corresponding behaviour for the s-channel c.m. M functions (by this we mean the M function with as arguments the s-channel c.m. momenta) which is given by

$$\begin{array}{ll} M & (p & c.m.) & < \left[\sin \theta \right]^{\frac{1}{2} \left[n + b + c + d \right]} \end{array} \tag{2.1.2}$$

(2.1.2) may be proved as follows:

The principle (P) enables us to determine the behaviour of $M(p_n)$ at the points $s = (m_i \pm m_i)^2$ by transforming to a frame without this singularity. The singularity at the physical region boundary is however present in all frames but as $\sin \theta_- \longrightarrow 0$ the c.m. vectors become invariant under rotations about the z-direction. So using the Lorentz transformation law (2.A.6) plus the principle (P) to enable us to expand N in powers of the x and y components of particles 3 and 4. immediately gives the result (2.1.2).

Using the principle (p'), the Trueman-Mick crossing relation (2.2.8) emables us to expose the nature of the singularity of W_i at = 6n₂ = n₂ 1 in the 4 ² matrices, since (i), has no such singularity. By this means, Mara and Wang have derived simple linear combinations of W_i which are everywhere nonsingular. Oxponenating this advantage of simplicity, these monitouses suffer the disadvantage of containing kinematic zeros. Thus their manipticity, prevent from the TW crossing relation (2.2.8), is insufficient to enable us to invert (2.2.8) and prove the input information that W_i is non-singular at s-thresholds. We will make this in chapter 4 when we consider partial wave residues. The analogous method to that of Hars and Hung will, for instance, say correctly that D waves are analytic functions of a at thresholds, but not that they vanish like s = (n₁ - n₂)².

To Section 2.5, we show how hars and Wang's arguments may be extended to derive linear combinations without singularities or zeros (but more complicately related to \parallel_2). Unfortunately this argument only works directly if but one particle has spin when only one super X₁ appears in the TW crossing relation. An extension to more than one spinning particle has only proved possible by the use of M functions. Thus if only particle 1 has spin; \parallel_4 is escentially identical to the M function with particle 1 at reat and particle 2 along the z-axis, while \parallel_4 is similarly related to the M function if in particle 2 along the z-axis.

can be seen from (2.A.6), all particles transform with the same reastion, so the extension to greater than one spinning particles only requires the CG. series union using M functions. The invertent mapitudes obtained in Section 2.5 are identical with those given earlier by HmppH21 and Hilliams²². Thus this section is not intended to be directly useful, but is included to illustrate the complete equivalence of the approach bised on (P') and that on inversion amplitudes. The requirement of partic conservation is not simply expressible in terms of M functions, as is familiar from the 4 components necessary in the Dirac formalism to describe a particy conserving spin § theory. We conclude Section 2.3 with a brief discussion of how one may try to overcome this difficulty.

Finally in Section 2.4, we present a selection of the various types of multitudes one can choose from when techling a particular problem. Here we compare the analyticity structure of M functions; singer multitudes; ordinary, parkly-connerving ann perpendicular $^{3/3}$ helicity multitudes. Also we introduce for me once to compare the behaviour at s = $(n_1-m_2)^2$ with that at $(n_1+m_2)^2$.

2.2 Derivation of the Trueman-Wick Crossing Relation

. The purpose of this section is to derive the results of $Muzinich^{(M)}$ and Truesan and $Wick^{(T)}$ with sufficient attention to detail to obtain the phase in their relation. This is

largely academic except in the case of elastic reactions where crossing leads to the same reaction and the conventional phases cancel out.

The most elegant derivation is that of Trueman and Mick, who take that amplitudes that transform in the same way under a general Lorentz transformation are identical up to a phase. To be pedantic this is not true as, for instance, an amplitude and its parity conjugate transform in the same way but are not necessarily equal. Indeed in a rather careless application of this method hislas and svenseon^{8,1} manage to "prove" heraltean analyticity. Such a method is presumably unable to determine the phase. Mazinich, by using the M function formalism, has however introduced a method which is sufficient for our purpose hore.

The calculation is rather tedious and so we simply outline the main steps and give the final result in (2,2.8).

We wish to determine the relation between the analytic continuation of H_8 and H_1 in the t physical region. From (2.A.3) and (2.A.7) we have the relation

 $d_{\lambda_1\nu_1}^{s_5}(-\theta_s) d_{-\lambda_2\nu_2}^{s_4}(-\theta_s) M_{s\nu_1\nu_2}(\lambda_2-\lambda_1) e_s^{c,m}$

(2.2.1)

 H_t is similarly expressed in terms of $M_t(p_t^{c.m.})$. M_s and M_t are essentially the same and the required factors in their relation is given in (2.4.8).

pc.m. = A pc.m

If we can find a Lorentz transformation A such that

$$p_{1,0}^{C,m} = A p_{1,0}^{C,m}$$

$$p_{2,0}^{C,m} = -A p_{2,0}^{C,m}$$

$$p_{2,0}^{C,m} = -A p_{2,0}^{C,m}$$
(2.2.5)

then we have

$$\begin{split} & \overset{\text{H}}{s}^{\lambda_{j}\lambda_{k}^{\prime}} i_{\lambda_{2}\lambda_{1}^{\prime}} = \overset{\text{d}}{a} \underset{\lambda_{j} \mid \underline{t}}{\sum_{l}} \left[\overset{\text{d}}{b} \cdot \overset{\text{d}}{b} \overset{\text{d}}{a} \cdot (\overset{\text{d}}{a} \cdot \overset{\text{d}}{b}) \overset{\text{d}}{a} \overset{\text{d}}{b} \cdot (\overset{\text{d}}{b} \cdot \overset{\text{d}}{a} \cdot \overset{\text{d}}{b} \cdot (\overset{\text{d}}{a} \cdot \overset{\text{d}}{b} \cdot \overset{\text{d}}{b} \cdot \overset{\text{d}}{a} \cdot \overset{\text{d}}{b} \cdot (\overset{\text{d}}{a} \cdot \overset{\text{d}}{b} \cdot \overset{\text{d}}{b} \cdot \overset{\text{d}}{a} \cdot \overset{\text{d}}{b} \cdot \overset{\text{d}}{a} \cdot \overset{\text{d$$

where 1st contains the conventional crossing phases, "particle
2° phases J1) and particle re-ordering phases. It is therefore

pases II and particle re-ordering phases. It is therefore
$$\int_{0}^{\infty} st = \epsilon_{23} \epsilon_{34} \epsilon_{42} \lambda_{3}^{*} \lambda_{2} (-1)^{s_{2} - \lambda_{2} + s_{4} - \lambda_{4}} (2.$$

(2.2.5)

(2.2.6)

The resistions in (1.2.2) are (almost) Migner rotations and it is clearly desirable to use the geometric method introduced by Mick¹⁰⁰) to evaluate them. In order to obtain the desired precision we first reduce A to a real Lorentz transformation, after which we may use geometry with a greater hope of getting the correct sign. We will treat particles 1 and 5 free and consider 1 and 6 later.

where ϕ_{λ} is a trivial combination of a complex boost in the z-direction with parameter $\theta^{\mu}=\pm 1\frac{M_{\chi}}{2}$ and realising abrough we about the y and z axes. The exact value of Λ_{λ} depends on the continuation alogated for Π_{λ} in travelling from the s to t physical region. Λ_{λ} is a fixed real Lorentz transformation such that $\Lambda_{\lambda}^{10} \hat{\Gamma}_{\lambda}^{10}$, is the set of vectors

A is a complex Lorentz transformation. We write A = A,A,,

where $\cosh |\Theta_{g}| = -\cos \Theta_{g}$ is $\geqslant 1$.

This is the set of vectors we would naturally have introduced if what considered decomposing the scattering amplitude wir.t. the 2+1 Lorents group for a:0, as opposed to the J dissensional restation group appropriate in the a-channel physical region. As is necessary for a negative mass squared representation the little vector $q_1 - q_2 = q_4 - q_3$ is $c \in (0, 0, 1, 0)$.

We split up the rotations in (2.2.4) as, for instance,

$$d \bigg[\mathsf{Cb}^{-1}(\mathsf{p}_{2n}^{\mathsf{c.m.}}) \mathsf{Ab}(\mathsf{p}_{2t}^{\mathsf{c.m.}}) \bigg] \ = \ d \bigg[\mathsf{Cb}^{-1}(\mathsf{p}_{2n}^{\mathsf{c.m.}}) \mathsf{A}_2 \mathsf{b}(\mathsf{q}_2) \mathsf{b}^{-1}(\mathsf{q}_2) \mathsf{A}_1 \mathsf{b}(\mathsf{p}_{2t}^{\mathsf{c.m.}}) \bigg]$$

Now both λ_2 and λ_1 have simple Higner rotations for particles 1 and 2: that of λ_1 may be evaluated geometrically to give the $d^{\frac{1}{2}} \lambda_1^2$ (see figure 2.1 and Mack⁹⁵) for the rules) and that of λ_2 algebraically. We have now evaluated half of the excression (2.2.4).

To consider particles 3 and 4 we use a frame similar to (2.2.6) but with 34 rather than 12 along the z-axis. Thus we write

$$A = A_2 X^{-1} | (x_{|Q_3|} | A_1)$$
 (2.2.7)

where $X_{\left[\frac{Q}{8}\right]}$ is a boost in the x-direction with parameter $\left[\frac{Q}{8}\right]$ given after (2.2.6). In figure 1, $X_{\left[\frac{Q}{8}\right]}$ A_1 takes us from C_t to X', while X'X represents the pure Lorentz transformation

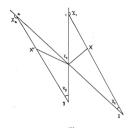


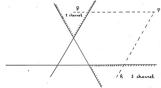
Figure 2.1: Momentum diagram 13):

 $\overline{\chi}_{i}$ are defined in (2.A.1d); C_{t} is the t-channel centre of mass frame; X is the frame (2.2.6); X' is the similar frame with 34 replacing 12; X and X' lie, as drawn, between the particle vectors for s < - max $\left\{ \begin{vmatrix} m_1^2 - m_2^2 \end{vmatrix} \right\}$ $\left\{ \begin{vmatrix} m_1^2 - m_2^2 \end{vmatrix} \right\}$

 $X_{\left|O_{S}\right|}$. Figure 2.1 is accurate for s,t large and near u=0, and by going to the physical region boundary the phases in relations such as $\{2,2,7\}$ may be evaluated.

The $d^{T_k}(\Theta_g)$, in the s-channel boosts $b(p_{3,ds}^{C,m})$, cancels with the $X_{[p_n]}^{T_k}$ to ensure that $A_2X_{[p_n]}^{T_k}$ has a simple Higner rotation for 3 and 4, while as before $X_{[Q_n]}$ A_1 give us the $d^{T_k}(X,)$.

Thereby one obtains the T.W. crossing relation with a phase that depends on the route taken from the s to the t channel. Rather than give the general case we state the answer for the continuation used in Regge theory.



<u>Figure 2.2</u>: Route used for H_g in TW crossing relation. H_g is defined at R, continued at fixed s to P, and then continued at fixed large t to Q.

(2.2.8)

In the continuation, singularities are met at both the s and t physical region boundaries, and $s = (m_1 \pm m_2)^2$, and

 $(m_3 \pm m_4)^2$.

The route adopted around the physical region boundaries

is such that from $R \to P \not = \frac{1}{2} + i \not = \frac{1}{2}$; and from $P \to Q$ $|\phi|^{\frac{1}{2}} \to -i |\phi|^{\frac{1}{2}}, \text{ while the route around } s = (m_1 + m_2)^2 \text{ is}$

Lemples 5 plane



and similarly for $s = (m_3 \pm m_4)^2$.

These conventions are consistent when in equal mass cases the singularities coincide.

Then finally the TW crossing relation is

$$H_{a}^{\lambda_{3} \lambda_{4}; \lambda_{5} \lambda_{6}} = \epsilon_{13} \epsilon_{34} \epsilon_{44}^{2} \lambda_{5}^{2} \lambda_{6}^{2}$$

$$d_{\lambda_{5} \mu_{5}}(x_{5}) d_{\lambda_{5} \mu_{6}}(x_{6}) d_{\lambda_{5} \mu_{5}}(x_{5}) d_{\lambda_{5} \mu_{6}}(x_{6})$$

6 - r. m. [2" - 2" + h" + h" + h" + h"] H" hrh+; h" h

The result for other routes of continuation may be derived

2.3 Derivation of Invariant Amplitudes

a) We take the TW crossing relation for a single particle (no.1) With spin. As pointed out in Section 2.1, the helicity supplictudes H₀ and H₀ are then essentially identical to the M functions, with particle 1 at rest and, respectively, particles 2 and 3 slong the 1-axis. In fact, independently of the TW crossing relation, it is obvious from figure 2.1 that TT-X₀.

is the rotation relating these rest frames.

We can trivially extend the treatment to any number of spinning particles by using the C.G. series. Let (j,m) label an irreducible state so formed. We have

$$M^{(j,m)}(p_{s}^{rest}) = (-1)^{m} M^{(j,p)}(p_{t}^{rest}) d_{pm}(X_{j})$$
 (2.3.1)

in the t physical region for the route of continuation such that $s_{1,2} \to [s_{1,2}]$, $\phi^{\frac{1}{2}} \to [\phi^{\frac{1}{2}}]^{\frac{1}{2}}$.

As described in Section 2.1 our statement of analyticity

- (i) At the physical region boundary M(prest t) is proportional to 4 imil times a function regular there.
- to $\phi^{2^{-m}}$ times a function regular there. (ii) While the other singularities in this case are: $M(p_g^{rest})$ is singular at $S_{12} = 0$; $M(p_g^{rest})$ is singular at $T_{13} = 0$.

We will now try to find some nonsingular linear combinations of $M(p_{\rm res}^{\rm res}t)$, which are not only analytic at $S_{1,2}=0$ but whose analyticity there enables one to deduce that $M(p_{\rm res}^{\rm res}t)$ is

analytic there. These linear combinations should also enable one to prove the physical region boundary behaviour of M and show that $M(p_{\perp}^{rest})$ is analytic at $T_{1,r} = 0$.

This problem is fortunately soluble in triangular fashion.

Thus we consider the equation: for any θ and $\eta = \pm 1$

$$d_{jm}^{j}(\theta) + \eta d_{-jm}^{j}(\theta) = \left\{ d_{j\gamma}^{j}(\theta - x_{i}) + \eta d_{-j\gamma}^{j}(\theta - x_{i}) \right\}$$

$$(2.5.2)$$

Then for

1) q = (-1) j:

Operate on both sides of (2,3.2) with

$$\frac{\sin^{(j-1)}x}{\xi!}, \quad \frac{d^{\frac{1}{2}}}{d^{\frac{1}{2}}(\omega + 0)} \quad \frac{1}{\sin^{j}\theta} \quad \dots \quad \left| \begin{array}{c} \text{for } \xi = 0 \dots \\ 0 = X \end{array} \right|$$

and write the resulting equation

We have

A2m = (-1)3-m B3m

(2.3.4)

(2.3.5)

Now form

$$\sum_{i=1}^{m} \theta_{i,m}^{g} (-i)_{i,m} W_{(i^{2},m)} (b_{i}^{2} e_{ig}) = \sum_{i=1}^{m} \theta_{i,j}^{g} W_{(i^{2},k)} (b_{i}^{2} e_{ig})$$

where $\boldsymbol{\ell}$ runs from 0 to j. Then from (2.3.4) we have a triangular structure with linear combinations of $M(\rho_{g}^{rest})$ (of one less in humber each time) equal to linear combinations of $M(\rho_{g}^{rest})$ (of one more in number each time).

We must now multiply by some factors to remove the singularities, and so we form the invariant amplitudes:

$$\mathbf{I}_{\frac{1}{6}}^{J} = \sum_{n \geq 0} \widehat{\mathbf{H}}_{\frac{1}{6}}^{J,n} \cdot \omega_{n}^{J} \left\{ \widehat{\mathbf{H}}_{1}^{J,n-n}(\mathbf{p}_{1}^{+m}) + (\omega_{n}^{J,n}) \widehat{\mathbf{H}}_{1}^{J,n-n}(\mathbf{p}_{1}^{+m}) \right\} \cdot (2.5.6)$$

$$\frac{\left[\times \mathbb{1}_{\frac{1}{6} - \frac{1}{6} -$$

These are proved analytic at $T_{13}=0$ from (2.3.6), and at $S_{12}=0$ from the expression (2.3.5) in terms of $M(p_t^{rest})$, and at $\phi=0$ from either (2.3.5) or (2.3.6).

But as they are triangular those relations may be inverted. Taking for example the $T_{1,2}=0$ singularity we start at $n=\ell=j$ and work dommurds in m, inverting to find $\mathbf{M}^{1,2}(Q_{j}^{\text{rest}})$ in terms of $\mathbf{T}_{ij}^{\text{rest}}$ ($\mathbf{X}>\mathbf{n}$). At each stage the new $\mathbf{M}^{1,2}(Q_{j}^{\text{rest}})$ has a simple coefficient in $\mathbf{T}_{i=m}^{\text{rest}}$ proportional to (sin $X_{1}^{\text{rest}})^{1-m}$, the $T_{1,2}=0$ o singularity of which is cancelled by the division factors in (2.7.6). Thus we may prove $\mathbf{M}(P_{j}^{\text{rest}})$ analytic at $T_{1,2}=0$. We can consider $T_{1,2}$ and j=0 similarly and hence find that T_{i}^{c} are the desired singularity free maplitudes.

II) $0 = -(-1)^{j}$ gives the remaining j invariant amplitudes. This time we operate on both sides of (2.3.2) with

$$\frac{\sin^{(j-\ell)}X_1}{(\ell-1)!} \quad \frac{d^{\ell-1}}{d^{\ell-1}(\cot\theta)} \quad \frac{1}{\sin^{j-1}\theta} \cdots \quad \bigg| \quad \theta = X_1$$

and write the resulting equation

$$\bar{X}_{\theta}^{jm} = \bar{b}_{\theta}^{j} \forall d_{pm}^{j} (X_{1}) \qquad (2.3.7)$$

for £ = 1 ... 1

where
$$\bar{A}_0^{jm} = (-1)^{j-m} \bar{b}_{2j-5-21}^{jm}$$
 (2.5.8)

Then, similarly to the previous method, we find the invariant amplitudes

$$\overline{\underline{\mathbf{I}}} \stackrel{J}{\underline{\mathbf{f}}} = \sum_{m \neq 0} \overline{\underline{\mathbf{h}}} \stackrel{J}{\underline{\mathbf{f}}} \stackrel{m}{\underline{\mathbf{f}}} \stackrel{(-)}{\underline{\mathbf{f}}} \frac{\left\{ \underline{\mathbf{m}}^{(j_1,m)} \left(\underline{\mathbf{p}}_{j_1}^{\text{start}} \right) - (-i)^m \underline{\mathbf{m}}^{(j_2,m)} \left(\underline{\mathbf{p}}_{j_1}^{\text{start}} \right) \right\}}{S_{j_1,j_2} \times \underline{\mathbf{f}}_{j_2}^{\text{start}} - \frac{1}{2} \overline{\underline{\mathbf{f}}} \frac{1}{\underline{\mathbf{f}}}} \frac{m^{(j_2,m)} (\underline{\mathbf{p}}_{j_1}^{\text{start}})}{\underline{\mathbf{f}}_{j_2}^{\text{start}} - \underline{\mathbf{f}}_{j_2}^{\text{start}}}$$
(2.3.9)

We can now compare these results with those from other methods. Firstly the method of Hara and Wang would form

$$H W_m^J(s) = \frac{M^{(j,m)}(p_s^{rest}) + (-1)^m M^{(j,m)}(p_s^{rest})}{S_{s_s^J}^{-1} \phi^{l_s^m}}$$
 for m > 0
(2.5.10

$$\mathcal{L}_{n}^{j} = \frac{\mathcal{L}_{n}^{(j,m)} \left(s_{j}^{\text{rest}} \right) - \left(-s_{j}^{m} \, \mathcal{N}^{(j,m)} \left(s_{j}^{\text{rest}} \right) \right)}{\mathcal{L}^{j-1} \, d^{m}} \qquad \qquad \text{for moo}$$

which are indeed nonsingular. We start off with Π_j^1 oc $\Pi^{ij}_{ij}(s)$, but in an endeavour to avoid kinematic zeros, we are forced to take in Π_{j-1}^1 a linear combination of $\Pi^{ij}_{ij}(s)$ and $\Pi^{ij}_{i-1}(s)$ and so on, eventually ending in Π_j^0 , which contains all $\Pi^{ij}_{ij}(s)$ and is in fact identical to $\Pi^{ij}_{ij}(t)$,

Secondly in our paper²¹² we have shown that the inveriant amplitudes 12. 12 are identical to those given certier by Hempel²² and milliams²⁰. We do not give the details here as it requires an excess of definitions and they do not appear to be very useful in the problems of interest in this thesis. Thus the Lorentz transformation in (2.A.6) of the N functions is the same whetever the number of spinning particles involved, but the constraint of parity conservation re-introduces a dependence on, for instance, the masses of the particles. In particular, it does not preserve the C.C. series that we conservation of parity implies

$$M(\underline{p}) = \mathbf{0} \int_{\underline{p}} D\left[\frac{\sigma_{+}\underline{p}}{m}\right] N(-\underline{p})$$
 (2.5.11)

or, restricting ourselves to the case when only two particles have spin,

$$P_{\lambda_1 \lambda_2}$$
 (p) = Φ γ_p $M_{\lambda_1 \lambda_2}$ (p),

where P is the parity conjugate N function defined by following

(2.3.11) with a rotation through π about the y axis to take - p back to p.

$$P_{\lambda_1 \lambda_2}(p) = (-1)^{s_1 + s_2 - \lambda_1 - \lambda_2} M_{-\lambda_1 - \lambda_2}(p) e^{-2 \lambda_2 \sigma}$$
 (2.5.12)

and $\sigma = \sigma_1^s + \sigma_2^s$, where σ_1^s and σ_2^s are defined in the appendix,

We may apply the principle (P) to $P_{\lambda_1\lambda_2}$ just as to M and thereby derive a new set of invariant amplitudes, which are the same linear combinations of P as I and I were of M. We must now choose a linearly independent subset of the new invariant amplitudes plus I and I for which parity takes a simple form. Fortunately this may also be done triangularly (I am only able to invert such matrices to show no singularities have been introduced!) and the results are given in our paper F1). For this special case - namely when only two particles have spin -Williams and Guertin have also derived invariant amplitudes. These seem more elegant and also achieve the discrete symmetry of particle identity if $s_1 = s_2$, which is not done by my method . However if $s_1 \neq s_2$ my method seems easier to invert. Thus although the general problem of obtaining parity-conserving invariant amplitudes for arbitrary spins has not been solved. in any problem of practical interest the initial writing down of a parity-conserving set is not difficult. The major portion of the work lies in obtaining both formulae for Hg in terms of

invariant amplitudes and inversely the invariant amplitudes in terms of ${\rm H}_{\rm g}$. Either direction is easy but having chosen one the other is difficult.

We sed this section with two disconnected comments. Firstly NeppU3 has shown that the non-singularity of the transformation (2,3-11) implies that it may always be diagonalized to find a singularity-free set of parity-connerving amplitudes. So our quest above to find their explicit form was not doomed to failure from the start. Secondly we have triest to apply the triangular method used in the beginning of this section to the full TW crossing with more than one angle. It produced an elegant set of invariant amplitudes for the case $s_1 = s_2 = 1, s_3 = s_4 = 0, \text{ or } \rho_3 = 1$ (which is soluble sums spins but $\Phi/\rho_3 = -1$ (even though this has freer independent amplitudes, i.e. four, against five in the former cases).

2.4 Comparison of the various AmplitudeFused for Spinning Particles

We describe and compare the properties of various types of amplitude that are found to be useful when considering particles with spin.

i) M functions

These are functions of momenta whereas the physical scattering amplitudes are properly regarded as functions of the boosts. They are therefore used in field theory, where integrals over all momenta occur. If we substitute for the argument $p = p_a^{C.m.}$, we find a function which is singular at s = 0; $s = (m_1 + m_2)^2$ when 1 and 2 are along the z axis; and on the physical region boundary as given in (2.1.2). The M functions behave in the same way at $s = (m_1 - m_2)^2$ as they do at $s = (m_1 + m_2)^2$ because the momenta have similar behaviour at those two points. This contrasts with the helicity and Wigner amplitudes which will be considered later. The singularity at $s = (m_1 + m_2)^2$ is easily exposed by rotating the vectors to put 3 and 4 along the z axis. If parity is conserved, the singularity at $s = (m_1 + m_2)^2$ is such that it may be removed by an overall factor in the manner of Hara and Wang, If $m_s \neq m_a$, the singularity at s = 0 is exhibited by taking particle 1 to rest.

ii) Wigner Amplitudes

These may, for instance, be found in Goldberger and Watson^{G1)} who justly claim them to be valid relativistically. Taking those corresponding to 1 and 2 along the z axis, they are related to helicity amplitudes by

$$w_{s:(12)} = \frac{\mu_3 \gamma_4^{1} \mu_2 \gamma_1}{\mu_3 \gamma_4^{1} \mu_2 \gamma_1} = \frac{u_4^{5}}{\mu_4^{5} \lambda_4^{5}} (\Theta_s) \frac{u_5^{5}}{\mu_3^{5} \lambda_3^{5}} (\Theta_s) \frac{u_5^{5} \lambda_3^{5} \mu_2^{5} \mu_1}{u_5^{5} \mu_2^{5} \mu_1^{5}}$$
(2.4.1)

They are singular at the physical region boundary, where they are proportional to sin 9 1/2 | Yi-Yr| P. = P. + Po. P. = P. + P.). The other singularities are at s = 0, which is not easily dealt with, and at the three thresholds $s = (m_1 \pm m_2)^2$ and $s = (m_3 - m_Z)^2$. If parity is conserved these latter singularities may be removed by an overall factor. Wigner amplitudes behave nonrelativistically like M functions. Thus comparing (2.2.1) and (2.4.1) we see the only essential difference between $M(p_8^{C.B.})$ and $W_{g_1(12)}$ lies in the boosts $e^{\sigma_1^2}$. These tend to 1 at s = $(m_3 + m_4)^2$, but at $s = (m_x - m_z)^2$, for instance, $sh \sigma_3^s \rightarrow 0$ but $ch \sigma_3^s \rightarrow 1$ or -1, if m. > or < m, respectively. This mass dependence and difference between $(m_3 - m_2)^2$ and $(m_3 + m_2)^2$ is characteristic of physical scattering amplitudes, as it is a singularity of the boost not the momenta. The e associated with M functions give them the same behaviour at $s = (n_{\pi} \pm m_{\chi})^2$ but destroy the parity transformation law. This is another general feature: the tedious calculations associated with form factors have little effect except schieving the correct behaviour at both s = (m. ± m,)2.

Wigner amplitudes are usually expanded in terms of orbital

$$\begin{split} & \bigvee_{\substack{\xi_1, (i_2)}} p_1 p_2, p_1, p_2, p_2 &= \sum_{\xi_1, \xi_2} C \left(\xi_1, \xi_2, \xi_1, p_2 \right) \cdot C \left(\xi_1 \xi_2, \xi_2, p_2 p_2 \right) \\ & \sum_{\xi_1, \xi_2} \left(\xi_1 \xi_1, p_1 \right)^{\xi_1} \cdot C \xi_2 \xi_2, p_1 \xi_2 - \left(\xi_2, \xi_2, T + p_1 - p_2, p_2 \right) \\ & C \left(\xi_1, \xi_1, T + 0, p_1 \right) \cdot - \frac{T}{T}, \xi_2, \xi_1, \xi_2 \\ & \sum_{\xi_1, \xi_2} \left(\xi_1, \xi_2, \xi_3, \xi_3, \xi_4 \right) \cdot \frac{T}{T}, \xi_3, \xi_4, \xi_4 \\ & \sum_{\xi_1, \xi_2} \left(\xi_1, \xi_2, \xi_3, \xi_3, \xi_4 \right) \cdot \left(\xi_2, \xi_3, \xi_4, \xi_4 \right) \cdot \left(\xi_3, \xi_4, \xi_4 \right) \cdot \left(\xi_4, \xi_4 \right) \cdot \left(\xi_4, \xi_4, \xi_4 \right) \cdot \left(\xi_4, \xi_4 \right$$

from which we find as usual $\frac{\tau_{I_1 I_2}^2}{t_1} \exp \left[s - (s_1 + s_2)^2\right]^{\frac{1}{2}} \frac{I_1}{t}$. $\left[s - (s_1 + s_2)^2\right]^{\frac{1}{2}} \frac{I_1}{t}$ but that this behaviour is not in general achieved at $s = (s_1 - s_2)^2$. To discuss the behaviour there we define.

iii) W' and H' Amplitudes

An discussed in ii) the only difference in behaviour of helicity amplitudes at $(n_1-n_j)^2$ and at $(n_1+n_j)^2$ lies in the $e^{i\theta}$ factors in (2.2.1). We thus define quantities to be identical with these factors at $(n_1-n_j)^2$ rather than at $(n_1-n_j)^2$. Namely put

$$H_s^i \lambda_3 \lambda_4 : \lambda_2 \lambda_1 = e^{i\pi(\lambda_{lighter}^i - \lambda_{lighter}^f)} H_s^{\lambda_3 \lambda_4 : \lambda_2 \lambda_1}$$
 (2.4.3)

where, for instance, $\lambda_{1ighter}^i = \lambda_k$ where k is the lighter particle of 1 and 2.

Similarly define H' and $T_{t_1}^{1:t_1^{t_2}}$ by (2.4.1) and (2.4.2) with primes added to all quantities. The $T_{t_1^{t_2}}^{t_3}$ are in general

completely different from τ_1^T and behave like $\left[n - (n_1 - n_2)^2 \right]^{\frac{1}{2}} \tilde{\chi}_1^T \sum_{i=1}^{T} \left[n_i - n_2 \right]^2 \right]^{\frac{1}{2}} \tilde{\chi}_1^T$. If parity is conserved one can however say that the singularity at the lower thresholds in τ_{i+1}^T may be removed by an overall factor. Notice that for the prison states the effective parity of the lighter particle of 1 and 2 is changed by (-1) $\frac{2n^2}{4}$ ighter. This has to be borne in mind when calculating the allowed values of t^2 . It is illustrated in Obstact t where we consider

Notice, as a particular case of the difference between prised and unprised states, that the total spin s_r is not equal to s_r, showing there is no significance of s_r eway from the physical region. Thus acting on H or W, s_r states are formed from

and s' states from

the NN vertex.

This also occurs at s=0 in equal mass cases and has been remarked upon by Freedman and Wang^{F2}).

iv) Helicity Amplitudes J1)

These are the most generally useful amplitudes, whose

analyticity has been given in the principle (F*). They are also singular at all the thresholds and this may be exposed, either by the TW crossing relation, or more easily by the relation (2.2.1) to M functions, with either 12 or 5% along the \$ xats. If \$a_{20}\$ and \$a_{20}\$, they are nonsingular at \$ = 0, which is useful advantage over N or W functions. However the latter are similar at the thresholds.

v) Purity-conserving Helicity Amplitudes
Hara H1) and wang W1) form

$$\mathsf{HW}_{\mathfrak{s}:(\eta_{\mathfrak{p}})}^{\lambda_{\mathfrak{f}}\lambda_{\mathfrak{q}}:\lambda_{\mathfrak{p}}\lambda_{\mathfrak{q}}}(\mathfrak{s},\mathfrak{t}) = \tilde{\mathfrak{s}}_{\mathfrak{s}}^{\lambda_{\mathfrak{f}}\lambda_{\mathfrak{q}}:\lambda_{\mathfrak{p}}\lambda_{\mathfrak{q}}}(\mathfrak{s},\mathfrak{t}) + \eta_{\mathfrak{f}} \tilde{\mathfrak{s}}_{\mathfrak{s}}^{-\lambda_{\mathfrak{f}}-\lambda_{\mathfrak{f}}:\lambda_{\mathfrak{p}}\lambda_{\mathfrak{q}}}(\mathfrak{s},\mathfrak{t}) \quad (2.4.4)$$

where

$$H_s = \frac{H_s}{B_s^{\lambda_{\Gamma} \lambda_{\dot{\lambda}}}}$$
,

and the physical region boundary function is given by

$$\theta_g^{\lambda_f \lambda_{\bar{i}}} = \left[\frac{1 + \cos \theta_f}{2}\right]^{\bar{i}_{\bar{i}}^{\bar{i}} \bar{\lambda}_{\bar{i}} + \bar{\lambda}_{\bar{j}} \bar{i}} \left[\frac{1 - \cos \theta_f}{2}\right]^{\bar{i}_{\bar{i}}^{\bar{i}} |\lambda_{\bar{i}} - \lambda_{\bar{f}}|}$$
 (2.4.5)

It is important to note that "parity-conserving" is not used in the same sense as in our Section 2.5 on invariant amplitudes and here means that a Regge pole of definite ∇P contributes to the amplitude. ($\nabla = \text{signature occurs because plain parity would require } z \rightarrow - z$). The singularities of

these asplitudes may be removed by overall factors as described by Imag⁶¹), whose results are correct except in the Boson-Feraion case, where she takes the fermion heavier than the boson. (See 2.4 (ii) and (iii) for this wass dependence), These results are not obviously useful because of the kinematic seron both at thresholds and at a = 0, where the original helicity amplitudes were monsingular. They are certainly not useful for Regge theory as it is much easier and more accurate to study the partial save resides right that for full socilities.

The kinematic zeros at thresholds may be removed by the . use of percendicular helicity amolitudes.

vi) Perpendicular Helicity (or Migner) Amplitudes These were introduced by Kotanski^{Kl)} and correspond to

quantizing the spin along the y not the z axis. They are related to ordinary helicity amplitudes by:

$$\begin{split} p_{q}^{2} p^{2} 4^{10} 2^{10} 1(s,t) &= X_{p_{3} h_{3}} (\eta_{2}^{\prime}) \ X_{p_{4} - \lambda_{6}} (s_{2}^{\prime}) \ H_{p}^{\lambda_{3} \lambda_{4}^{\prime} \lambda_{2}^{\prime} \lambda_{1}} (s,t) \\ & X_{-\lambda_{2} p_{2}} \ (-\eta_{2}^{\prime}) \ X_{\lambda_{1} p_{1}} \ (-\eta_{2}^{\prime}) \end{split} \tag{2.4.6}$$

where $X(\Theta)$ is a rotation through Θ about the x axis and $X(\pi_{\sum}')$ carries the y into the z axis.

The TW crossing felation becomes diagonal when expressed in terms of these amplitudes. This means the non-singularity at s thresholds of $P_{\rm p}$ is easily expressed without kinematic

zeros in terms of P_a .

Unfortunately this is countered by the grave disadvantage that P_g and $P_{\bar{g}}$ are both singular at the physical region boundary. Thus it seems that they are only useful locally for expressing the threshold conditions on H_g in terms of simpler linear combinations than the original TW crossing relation and H_g gave.

We finish by listing the behaviour of P₈ at the thresholds in more detail. This requires a separate discussion for the three mass types. In Leader's notation, these are denoted,

$$n_1 \neq n_2$$
, $n_3 \neq n_4$ UU

 $n_1 = n_2$, $n_3 \neq n_4$ $\pi_1 U_\Gamma$ (2.4.7)

a) <u>U U</u>

We will need a phase to describe the route taken round the physical region boundary singularity to reach the thresholds, and suppose tan \mathbf{e}_a > i.X , X = 1 at thresholds. Then at the top thresholds s = $(\mathbf{e}_i + \mathbf{e}_j)^2$

$$P_s^{p_3p_4:p_2p_1} \sim s_{12}^{X(p_1+p_2)} s_{34}^{X(p_3+p_4)}$$
 (2.4.8a)

while at the bottom thresholds the result is the same, except the index of the lighter particle in each pair (12), (34) is reversed in sign.

At s = 0, P is nonsingular.

At s = 0 we have

b) 8, U The results at $(m_3 \pm m_4)^2$ and $(m_1 + m_2)^2$ are as in (a).

$$p_{p_{3}}^{p_{3}p_{4}:p_{2}p_{1}} \sim (\sqrt{s})^{\epsilon(p_{1}-p_{2})},$$
 (2.4.8b)

where the new phase
$$\epsilon$$
 is defined by
$$\phi \stackrel{i_2}{\to} n_1 \ (n_3^2 - n_2^2) \ i \, \epsilon \qquad \text{as } s \to 0 \ .$$

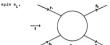
g g c) The result at the top thresholds is as in (a). s = 0 is now a physical region boundary and Pg are no longer wellbehaved there. However the TW crossing angles + TT, and so one can use d $(\pi/2)$ rather than X $(\pi/2)$ to express conditions on Hg, i.e. we have (independent of the route of continuation

of Hg)

 $\mathbf{d}_{\mathsf{p}_{3}\lambda_{3}}(\neg y_{2})\cdot \mathbf{d}_{\mathsf{p}_{2}} - \lambda_{2}(\neg y_{2})\cdot \mathbf{b}_{8}^{\lambda_{3}\lambda_{2}:\lambda_{2}\lambda_{1}} \mathbf{d}_{-\lambda_{3}\mathsf{p}_{3}}(\neg y_{2})\cdot \mathbf{d}_{\lambda_{1}\mathsf{p}_{1}}(\neg y_{2})$

Appendix: Notation T2), S1)

i) Particles are labelled $i = 1 \dots 4$ and have mass m_i and



Mandelstam invariants are defined as usual

$$s = (p_1 + p_2)^2$$
 $t = (p_1 - p_3)^2$ $u = (p_1 - p_2)^2$

ii) notations are labelled by Duler angles κ , ρ , Y and written $r(\kappa$, ρ , Y). Their matrix elements $D^T(c_k, \rho, Y)$ are defines as in Jacob and $\operatorname{Mic} x^{(k)}$ and σ^I_{mr} (0) = σ^I_{mr} (0,0,0)). This sifters by $(-1)^{m-k}$ from the definition used in the useful reference. The complex J, let thus define our second type functions σ^I_{mr} (5 to be $(-1)^{m-m}$ times those in Gunson and Andersenh.)

Pure Lorentz transformations from rest up to ϱ are written $h(\varrho)$, and $h(\varrho)$ is a general boost up to ϱ .

Vector components are written in the order (x, y, z, t). iii) The kinematic definitions are NL

$$\beta_{11}^2 = \left[s - (m_1 + m_1)^2\right]$$
. $\left[s - (m_1 - m_1)^2\right]$ (ij) = (12) or (34) (2.A.1a)

(2.A.1b)

(2.A.1c)

(2.A.1d)

and similarly for t and u.

The c-m frames are defined as usual (see section (v) of the appendix), and the s-channel c.m. vectors are denoted ρ_{1}^{tot} . i. = 1....4 and we take ρ_{2}^{rest} to be the system generated from this by a pure Lorentz transformation reducing particle 1 to rest.

The c.m. scattering angle is given by

$$s_{12} s_{34} \cos \theta_s = s(t-u) + (n_1^2 - n_2^2) (n_3^2 - n_4^2)$$

 $s_{11} s_{34} \sin \theta_s = 2 \sqrt{s} \phi^{\gamma_3}$ dations ϕ

The boost parameters are defined by

ch
$$\sigma_i^s = E_i^s/n_i$$
,

where E_{i}^{s} are the c.m. energies.

The TW crossing angles are defined by

$$s_{12}\tau_{13}\cos \chi_1 = (s_{12}^2 - m_1^2 - m_2^2) (t_{12}^2 - m_3^2) + 2m_1^2 \Delta$$

$$\begin{split} s_{12}\tau_{24} &\cos \chi_2 = - (s + m_2^2 - m_1^2) (t + m_2^2 - m_4^2) + 2m_2^2 \Delta \\ s_{34}\tau_{13} &\cos \chi_3 = - (s + m_3^2 - m_4^2) (t + m_3^2 - m_1^2) + 2m_3^2 \Delta \end{split}$$

$$S_{34}T_{24} \cos X_4 = (s+m_4^2-m_3^2) (t+m_4^2-m_2^2) + 2m_4^2 \Delta$$

where $\Delta = m_{\pi}^2 + m_{\pi}^2 - m_{\ell}^2 - m_{\ell}^2$

iv) Single Particle States

Particles are described by invariantly normed states

 $|p, \lambda >$, which have Lorentz transformation law

 $U(\Lambda) \mid p, \lambda \rangle = \sum_{i=1}^{n} |\Lambda_{i}, \lambda' \rangle D_{\chi_{\lambda}} (\omega(\Lambda, p))$

for A & \$L(2, C) where the Wigner rotation

$$\omega (\Lambda, p) = b^{-1} (\Lambda_p) \Lambda b (p)$$

Here the boost b(p) is a Lorentz transformation taking us from rest to momentum p. Different choices of b(p)

correspond to the different type of spin amplitudes. For Helicity States: b(p)=h(p) $r(\psi,\,\varphi,\,-\psi)$

= r(φ, θ, -φ) h(|p| z)

Wigner States: b(p) = h(p)

states defined by

(2.A.3)

(2.A.2)

where ϱ is in direction (φ, φ) . Spinor states are obtained by extending D to be a representation of the homogeneous Lorentz group and splitting (2,A,2) into its constituent parts. We will only need lower spinor

 $| \underline{p}, \ll \rangle = \sum_{\lambda} |\underline{p}, \lambda \rangle D_{\lambda \infty} (C b^{T}(\underline{p}))$ (2.A.5)

which is independent of b. Thus we have simplified the Lorentz transformation law at the cost of a non-unitary transformation. Here

$$C = \begin{bmatrix} -1 \\ 1 \end{bmatrix} \in SL(2, \mathbb{C})$$
.

v) Scattering Amplitudes

As usual, define the T matrix by

We can take matrix elements of the T matrix between any of the above type of states. Choosing spinor states, we find M functions which have the Lorentz transformation law

$$M(p) = \Theta(\Lambda) M(\Lambda^{-1}p)$$
 (2.A.6)

This can be simplified by use of the C.G. series, thereby reducing the study of the M function describing many particles to that describing one.

The relation between M functions and physical scattering

where b are given by (2.A.3), (2.A.4) for Nigner or helicity amplitudes. It is also customary 21 1 to insert an extra factor $^{8}2^{-\lambda_{2}}(-1)^{8}4^{-\lambda_{2}}$ in (2.A.7) when dealing with helicity amplitudes.

In (2.A.7) we substitute the usual c.m. momenta for the arguments of M and T. In this thesis we shall always take the c.m. frame of $\left\langle \operatorname{cd} \mid T \mid \operatorname{ab} \right\rangle$ to be that with the positive

y axis along a x c, and with 0 6 0 4 T



We define the s-channel as <34 $|T_g|$ 21> and the t-channel as <24 $|T_b|$ |31>, where the ordering of labels defines both the momenta as above and the order of creation operators.

We will call Migner amplitudes $H^{V_4V_5V_2V_1}$ and helicity amplitudes $H^{\lambda_k\lambda_jt\,\lambda_2\lambda_1}$ adding a subscript s, t, ù to distinguish the various channels.

Also we put $\lambda_1 = \lambda_1 - \lambda_2$, $\lambda_f = \lambda_3 - \lambda_4$.

vi) Crossing

In terms of M functions this means we must add a label s, t, u to M in (2.A.7) and introduce a crossing phase λ_n , so that $|a>+\lambda_n<\bar{s}|$ on crossing, where \bar{s} is the antiparticle of a.

. λ_{-a} corresponds to the relative phase of the particle

annihilation operator s_{∞} and anti-particle creation operator b_{∞}^{\dagger} in the field theory approach of Weinberg $^{W_{\bullet}}$) and Carruthers C1).

where \ll is a spinor index. This is useful if one wishes to use Carruther's^{Cl)} SU₂ crossing relations.

Then, taking account of the change in the ordering of states, we have, on crossing 2 and 3,

$$M_{t}(p_{1},p_{2},p_{3},p_{4}) = \epsilon_{23} \epsilon_{34} \epsilon_{42} \lambda_{2}^{*} \lambda_{3} M_{s}(p_{1},-p_{2},-p_{3},p_{4})$$
 (2.A.8)

where $\epsilon_{ij} = +1$, unless i,j are both fermions when it is -1.

When we leave the subscript s, t, u off M, we will mean

vii) The Crossing Phase λ

Let γ_P γ_C γ_T be the phase factors in the transformations of the single particle states under the discrete transformations of parity, charge conjugation and time reversal respectively.

That is. for Wigner states:

$$P \mid \varrho, \lambda \rangle = \eta_{P} \mid -\varrho, \lambda \rangle$$

$$T \mid \varrho, \lambda \rangle = \eta_{T} \mid -\varrho, \lambda \rangle \triangleleft_{\chi_{\lambda}} (\tau)$$

$$C \mid \bullet \rangle = \eta_{C} \mid \bar{\bullet} \rangle$$

Then $\lambda_a = \lambda_{\overline{a}}$, and the PCT theorem says

 $\lambda_a = \int_P \int_C \int_T$ for bosons $\lambda_{\bar{a}} = \int_P \int_C \int_T$ for fermions

ignoring possible superselection rules. The I occurs for fermions because the physical PCT symmetry differs from the crossing relation known as the "PCT theorem" by a complete reversal in the order of all states in the scattering amplitude. Finally we note that we will take $\theta_0 = 1$ in this them;

CHAPTER 3

Partial Wave Amplitudes, Pole Residues and Form Factors

3.1 Introduction

In this chapter we describe some applications of the

Section 3.2 is devoted to a discussion of the kinematic singularities of partial wave amplitudes. This is necessary if, for instance, one wishes to include the 8 1256 in an N/O calculation, and from the sphiless case we know how important it is to enforce threshold behaviour to obtain successful results in anormistance calculations.

Section 3.3 is devoted to pole residues. It is well known 30 that to evaluate one-particle exchange diagrams, instead of using a specific Lagrangian, it is swiftlen to put a pole in the crossed partial wave series. We spell this out showing how one may cope with the three pieces of information we think are contained in a phenomenological Lagrangian. These are analyticity, crossing for the 2 +1 amplitudes and positive definite hamiltoniam. These are all easily expressed in an helicity formalism.

Finally in our last physical j application we discuss form factors and the solution they offer for the behaviour at s=0.

We also evaluate the helicity amplitudes for some well known form factors $^{\mbox{\scriptsize J2})}$.

Chapter 4 on Regge theory will be an extension of Section 3.4 and the work in Section 3.2 on threshold conditions.

3.2 Partial Wave Amplitudes

We define our helicity partial waves by

and we deal with the two cases separately.

$$\tau_{\lambda_{3}\lambda_{4}:\lambda_{2}\lambda_{1}}^{J}(s) = \frac{1}{2} \int_{-1}^{+1} d(\cos \theta) \ d_{\lambda_{2}\lambda_{1}}^{J}(\theta_{3}) \ H_{s}^{\lambda_{3}\lambda_{4}:\lambda_{2}\lambda_{1}}(s, t)$$
 (5.2.1)

Kinematic singularities occur at thresholds and at s = 0

(i) The Thresholds

This is the easier case, and we have already shown in Chapter 2 how one can define orbital angular momentum states in terms of which there is an exact statement of the behaviour at s = $(n_1 + m_j)^2$. We also defined primed states that behave similarly at $(n_1 - m_j)^2$. Namely:

$$T = \frac{1}{k_1} \frac{f_1}{k_2} = \frac{(2 \cdot k_1 + 1)^{k_1} (2 \cdot k_2 + 1)^{k_2}}{(2 \cdot 2 + 1)} \sum_{\lambda_1 \cdot \lambda_2 \cdot \lambda_3} \lambda_4$$
 (5.2.2)

while $T = \begin{cases} J:s_1^*s_2^* \\ \vdots \end{cases}$ is given by the same equation with the extra phase $iw (\lambda_1^1, \dots, \lambda_n^n)$

i* (\lambda_lighter - \lambda_lighter)

Alternatively one can take "parity-conserving" linear combinations in the manner of Hara and Wang,

Thus form

$$J_1(q)$$
 $T_{\lambda_3\lambda_4;\lambda_2\lambda_1} = T_{\lambda_3\lambda_4;\lambda_2\lambda_1} + \eta T_{\lambda_3\lambda_4;-\lambda_2-\lambda_1}$

$$Then at $s = (n_1 + n_2)^2$ if$$

$$(-1)^{J-s_1-s_2} \eta = +1 : \tau^J \propto \left[s - (n_1+n_2)^2 \right]^{\frac{1}{2}} \min \text{minimum even } t_1$$

= -1 :
$$\tau^{J} \propto \left[s - (n_1 + n_2)^2\right]^{\frac{1}{2} \cdot \text{minimum odd } \ell_1}$$

where ℓ_1 runs from the $\min[0, J-s_1-s_2]$ to $J+s_1+s_2$. Some kinematic zeros can easily be removed by forming total spin linear combinations.

$$\sum_{\lambda_1-\lambda_2=\lambda_i} \, {\rm C}(s_1s_2s_i\!:\!\lambda_1,\!-\lambda_2) \, \, {\rm T}^{\rm J:\,(\eta)}_{\lambda_3\lambda_4\!:\,\lambda_2\lambda_1}$$

for which & , runs up from min[0, J - si] .

At $s = (s_3 + m_q)^2$ we have similar results with $q \rightarrow q$ times the product of intrinsic parities of the particles $(-1)^{8}1^{+8}2^{-8}3^{-8}4$.

At $s = (m_1 - m_2)^2$ we should replace γ by $\gamma(-1)^{2s_{11ghter}^2}$.

In boson fermion cases it is customary to regard the amplitude as a function of f_0 . In the results for $f_0 = -|a_1| 2n_2$, as from the argument leading to primed states, we pick up phase factors $e^{\frac{2k-n}{N}}k$ from the $e^{\frac{2k}{N}}$ in (2.2.1) whenever cher s=-1.

For example, if $m_1 > m_2$

LTS complex plane

The correct behaviour for \sqrt{s} < 0 may also be obtained from the MacDowell symmetry relation $^{(4)}$, $^{(5)}$. Only in simple cases such as $^{(5)}$ $^{(5)}$ $^{(5)}$ $^{(5)}$ is those conditions a simple account without following

satisfy all these conditions simultaneously without introducing kinematic zeros. A possible approach is to take ordinary orbital angular momentum states to ensure the correct behaviour at the top thresholds, removing the singularity at the bottom threshold with an overall factor, thereby introducing the usual kinematic zeros.

An alternative method is suggested by Regge theory and is rather surprisingly applicable to physical J, with an important proviso to be mentioned later. If one expands asymptotically the Trueman-Mick crossing relation, one finds.

$$\frac{G}{J_{P_1P_4;P_1P_4}} = \frac{e}{e} \frac{d_{P_1\lambda_1}(X_1^{e})}{d_{P_1}^{e}(\frac{1}{1-\lambda_1+1})\frac{T}{P_1}\frac{T}{\lambda_1}\frac{1}{\lambda_2}\frac{1}{\lambda_1}\frac{1}{N_1}} = \frac{T}{d_{P_1}^{e}(\frac{1}{1-\lambda_1+1})\frac{T}{P_1}(\frac{1}{1-\lambda_1+1})\frac{T}{P_1}(\frac{1}{1-\lambda_1+1})} \cdot \left[\frac{g}{d_{P_1}^{e}d_{P_2}^{e}}\right]^{J}$$
(5.2.4)

 $\sin X_1 = \frac{2n_1 \sqrt{-8}}{8}$

is nonsingular at s-thresholds. Here $\mathfrak{X}_1^{\bullet 0}$ is given by (2.A.1d) with t taken to $\bullet \circ$, e.g.

$$\cos X_1^{\infty} = \frac{s + m_1^2 - m_2^2}{s_{12}}$$

The $\, \Gamma \,$ functions come from the asymptotic coefficient of the $\, \mathrm{d}^{\, J} \,$ function.

It is perhaps not obvious that this is the full kinematic condition, but this may be proved by using the method of Kotanski $^{\rm K1}$) to write (3.2.4)

$$P_{Y_{3}Y_{4}:Y_{2}Y_{1}} \sim S_{34}^{J+Y_{3}+Y_{4}} S_{12}^{J+Y_{1}+Y_{2}}$$
 (5.2.5)

As usual the index of lighter particle is reversed at the lower threshold, and

$$P_{X_3Y_6:X_2Y_1}^J = \sum_{\lambda,\lambda,\lambda} T_{\lambda_3\lambda_6:\lambda_2\lambda_1}^J \qquad (5.2.6)$$

$$\mathbf{x}_{\mathbf{Y}_{d}^{-\lambda}_{d}}^{\mathbf{x}_{d}}(\pi/2) \mathbf{x}_{\mathbf{Y}_{3}^{-\lambda}_{3}}^{\mathbf{x}_{3}}(\pi/2) \mathbf{x}_{\mathbf{J}_{A_{f}}}^{j}(\pi/2) \mathbf{x}_{\mathbf{J}_{2}^{-j}}^{j}(-\pi/2) \mathbf{x}_{-\lambda_{2}^{-\lambda}_{2}^{-j}}^{\mathbf{x}_{2}^{-k}}(-\pi/2) \mathbf{x}_{-\lambda_{2}^{-\lambda}_{2}^{-k}}^{\mathbf{x}_{2}^{-k}}(-\pi/2) \mathbf{x}_{-\lambda_{2}^{-k}_{2}^{-k}}^{\mathbf{x}_{2}^{-k}}(\pi/2) \mathbf{x}_{-\lambda_{2}^{-k}_{2}^{-k}}^{\mathbf{x}_{2}^{-k}}(\pi/2) \mathbf{x}_{-\lambda_{2}^{-k}_{2}^{-k}_{2}^{-k}}^{\mathbf{x}_{2}^{-k}}(\pi/2) \mathbf{x}_{-\lambda_{2}^{-k}$$

(3.2.6 continued)

We can now easily prove the equivalence of the orbital angular momentum conditions and (3,2,5) by acting the C.G. coefficients in (3,2,2) on the rotations by π/a about the x axis in (3,2,6).

(3,2,4) are splendid states except that for j cmax($s_1 + s_2$, $s_3 + s_4$) the "nonsense conditions" imply that not all the states are linearly independent. There seems no easy way round this, and we will comment on it later in Regge theory, where these conditions occur when the trajectory passes through physical values of j and not for all s.

We end by remarking that in our analysis at $\mathbf{s} = (\mathbf{a}_1 - \mathbf{a}_j)^2$ we excluded the case - which occurs in elsetic $(\mathbf{e}_2, \mathbf{w} \mathbf{h})$ scattering - where the threshold condition is spblit by a singularity $\mathbf{u} = (\mathbf{a}_1 + \mathbf{n}_j)^2$ entering the region of integration. If this happens it is best to split the molitudes into two parts - the major part without the singularity to which our theory applies, and a small portion treated separately. (ii) $\mathbf{s} = 0$

' Finally we discuss the behaviour at s=0. According to Freedman and WangF3),F4), this is dominated by the leading Regge pole and one should remove a factor $s^{-\epsilon}$. We do not

believe this is always the correct approach to N/D equations. Writing a fixed s dispersion relation for $H_a^{3} \lambda_a^{1/2} \lambda_a^{1/2}$ namely

$$y_{S}^{\lambda_{2}\lambda_{2}} \lambda_{2} \lambda_{2}(s, t) = a_{S}^{\lambda_{1}\lambda_{2}}(s, t) \left\{ \frac{1}{w} \int_{-\frac{1}{2}-\epsilon}^{\frac{1}{2}-\epsilon} c \frac{\lambda_{2}\lambda_{2}}{\epsilon^{2}} \lambda_{2}^{\lambda_{2}} + \frac{1}{w} \int_{-\frac{1}{2}-\epsilon}^{\frac{1}{2}-\epsilon} a_{S}^{\lambda_{2}} \lambda_{2}^{\lambda_{2}} \lambda_{2}^{\lambda_{2}} \right\}$$
(5.2.7)

where $a_s^{\lambda_1 \lambda_1}$ is defined in (2.4.5), we write T^J in Froissart-Gribov form

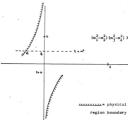
$$w = T_{\lambda_3 \lambda_2 \lambda_3 \lambda_3 \lambda_4}^J = -\int_0^\infty b_1^{\lambda_3 \lambda_2 \lambda_3 \lambda_3} e_{\lambda_2 \lambda_2}^{-1} (e_g(t')) de_g(t')$$

$$+ (-1)^{2 + \lambda_3} - \int_0^\infty b_1^{-1} \lambda_2^{-1} e_{\lambda_2^{-1} \lambda_2^{-1}} (\bar{e}_g(t')) d\bar{e}_g(t')$$
(5.2.8)

where $\bar{z}_{g}(u^{*}) = -z_{g}(En^{2} - s - u^{*})$ in the second integral while $e_{\lambda_{L}J_{L}}^{iJ} = e_{\lambda_{L}J_{L}}^{J}e_{g}^{\lambda_{L}J_{L}}(s)$ and e^{J} is defined in the appendix to Chauter 2.

We must now remove a function of s from T^J so that its left-hand cut discontinuity comes from the cut of e^t for z_L , $c_L = 1$ J rather than kinematic singularities of D_{c_L} or u^{t_L} . Not the case of UU scattering (see (2, 4, 7) for this

For the case of UU scattering (see (2.4.7) for this notation) the structure of the physical region near s = 0 is



. A \$-function contribution to \$\(\text{D}_{\text{s}} \text{ t = \$\text{n}^2\$ gives rise to,} \)
smoon other things, a left should cuit \$\text{n}^2\$ between \$A\$ and \$\text{N}\$. In particular if we take the discontinuity arising from a Regge pole for large t (and integrate up to t = \$\text{n}\$), we will find \$n\$ is \$^2\$ behaviour near \$\text{s} = 0\$. Hence, to enforce this behaviour is both unnecessary and an inferior approximation, so it replaces the cut from \$A\$ to \$\text{D}\$ you for from \$\text{s} = \$\text{m}\$\$ to \$\text{L}\$.

In the case of spin, the Regge contribution is not always the most important near s=0. We deal with the three mass types separately.

(3.2.9b)

(a) Thus in the UU case the Regge contribution to τ^2 is proportional to the residue β of the leading Regge pole, where the constant may be evaluated in the same way as $\tau^{2,3}$ using the Bateman manuscript $\tau^{2,3}$ and the hypergeometric form of e_{λ}^{1} (t_{λ}^{1}).

 $\begin{array}{ll} \lambda_1 \lambda_1 & & \\ \lambda_1 \lambda_2 & & \\ & \text{Putting } \epsilon = \text{sign } (m_1^2 - m_2^2)(m_3^2 - m_4^2), \ \beta \text{ must} \\ & &$

Meanwhile a general contribution from a finite part of the discontinuity $t=m^2$

$$\sim \frac{s}{c_1^{-1}\lambda_1^{-4}\lambda_1^{-1}} \left\{ \text{finite } + Is \left| \lambda_1^{-4}\lambda_1^{-1} \varrho_{n(\pm s)} \right| \right\}$$
 (5.2.9a)

It is this last form whith enables one to determine the factor $(F_a : \{\lambda_1 - \epsilon \lambda_1 | -2\})$ that multiplies τ^a , and ensures that the left hand cut comes only from $e^{i\lambda_a}_{i,k}(z)$.

(b) In the $U_i E_i$ case we are faced with a kinematic zero

(b) In the $U_r E_k$ case we are faced with a kinematic zero difficulty. In terms of the amplitudes of (5.2.3) we have $(s_1 = s_2)$

$$T_{\lambda_{j}\lambda_{k};\lambda_{j}\lambda_{k}}^{J:(\eta)} \sim \sum_{i=1}^{-s_{1}} \text{ if } \eta(-1)^{J+\lambda_{i}+\lambda_{j}+\lambda_{k}} = 1$$

$$f_{\overline{s}\cdot s}^{-s_{1}} \text{ if } \eta(-1)^{J+\lambda_{i}+\lambda_{j}+\lambda_{k}} = -1$$

It is possible to do better than this by defining yet

another type of orbital angular momentum state, by taking (2.2.1)

and replacing the boosts for particles 1 and 2 by their values at s = 0. Whortunately this produces an $e^{\frac{1}{2} \frac{N_1}{N_1}}$ and gives difficulty with partity, so I will not pursue it here. (c) In the Ef case it is well known from the NN example $^{(2)}$ that the conditions (2.4.8c) on \mathbb{H}_g require relations between the partial wave amplitudes of different J. The relation $^{(2)}$ (2.4.8c) experise on forming total pairs' states (section 2.4).

 $s_i^* = s_i^* = 0$: no condition

and in the NN case we have:

However if $\mathbf{s}_{\mathbf{f}}^1 = \mathbf{s}_{\mathbf{f}}^1 = 1$ all amplitudes are constant at $\mathbf{s} = 0$ but satisfy

$$\begin{array}{lll} -\sqrt{2\,(T+1)} \;\; {\boldsymbol{R}}^{T-1} \;\; + \; (T+1) \sqrt{\frac{2-1}{2}} \left({\boldsymbol{C}}^{T-1} - {\boldsymbol{B}}^{T-1} \right) \;\; + \; \frac{1}{2T+1} \left({\boldsymbol{B}}^T + {\boldsymbol{C}}^T \right) \\ + \sqrt{T\,(T+1)} \;\; {\boldsymbol{R}}^{T+1} \;\; + \; (T+1) \sqrt{\frac{2-1}{2}} \left({\boldsymbol{B}}^{T+1} - {\boldsymbol{C}}^{T+1} \right) \;\; \sim \;\; {\boldsymbol{S}} \end{array}$$

where
$$A^J = T^J \stackrel{1'}{0} \stackrel{1'}{0} \stackrel{1'}{0} = f_0^J$$

$$a^{J} = \tau^{J} \quad {1 \atop 1}^{1} \quad {1 \atop -1}^{1} \quad = \quad \frac{1}{2} \quad (r_{1}^{J} - r_{22}^{J})$$
 (5.2.10)

in more usual (and better!) notation.

 $C^{J} = T^{J} \begin{pmatrix} 1 & 1 \\ 1 & 1 \end{pmatrix} = \frac{1}{2} (\Gamma_{1}^{J} + \Gamma_{22}^{J})$

$$\sum_{\lambda_{f}\lambda_{\underline{i}}} \operatorname{d}_{\rho_{f}\lambda_{f}}^{s_{f}^{f}} (-\overline{\nu}/2) \overset{\lambda_{f}^{f}:\lambda_{\underline{i}}}{\operatorname{d}} \operatorname{d}_{\lambda_{\underline{i}}\rho_{\underline{i}}}^{s_{\underline{i}}^{f}} (\overline{\nu}/2) \sim I_{\overline{s}}^{|\rho_{f}^{-\rho_{\underline{i}}}|}$$

where $H_{\alpha}^{\lambda_{\Gamma}; \lambda_{\underline{i}}} = \sum_{(2J+1)} T^{J} J_{1}^{\alpha_{\Gamma}' \alpha_{\underline{i}}'} d_{\lambda}^{J} (Q_{\alpha})$

I will content myself by noting a method of generating relations such as (3,2,10) in general. It will take no account of the special forms of (3,2,1) and simply constants in finding a set of functions $f^{\hat{k}}$, independent of $\lambda_{\hat{k}}$ and $\lambda_{\hat{k}^*}$, in terms of which we may expand $d_{\hat{k}}^{\hat{k}^*}$. This is a generalization of the method used $i_{\hat{k}^*}^{\hat{k}^*}$, and a suitable set of $f^{\hat{k}}$ is

rt = dx0 (0s) t= x ... o

$= \lambda_1 + \lambda_f$ even

where
$$x = (s_{1}^{*} + s_{1}^{*})/2$$
 if $s_{1}^{*} + s_{1}^{*}$ even

or $x = (s_1^i + s_1^i - 1)/2$ if $s_1^i + s_1^i$ odd

b) $y^{\dagger} + y^{\dagger}$

where $x = (s_1' + s_1' - 2)/2$ if $s_1' + s_1'$ even

$$x = (s_{i}' + s_{f}' - 1)/2$$
 if $s_{i}' + s_{f}'$ odd

This may be proved by repeated use of the C.G. series and gives relations like (3.2.10) involving in general 2x + 1 values of j.

3.3 Pole Residues

We consider in decreasing order of importance and complication: analyticity, crossing and the conditions following from a positive definite Hamiltonian.

(i) Analyticity

If T^{J} , as defined by (3.2.1), has a pole at $s=m^{2}$, then near there

$$H_{S} \sim \frac{(2J+1)}{m^{2}-s} \quad \chi_{\lambda_{4}\lambda_{5}}^{J} \quad \chi_{\lambda_{2}\lambda_{1}}^{J} \quad \sigma_{\lambda_{1}\lambda_{1}}^{J} \quad (e_{g})$$
 (5.5.1)

The residue at the pole in N₂ is uniquely (and simply) specified, but so usual there is an ambiguity in the off-mass shell continuation. In a dispersion approach this means there is no such thing as the continuation of a pole to an amplitude as it depends on the direction dispersed in and the assuad behaviour at infinity. One could put the pole in invariant amplitudes but so if no not know a generally useful form for these, I will only describe the method of Hera and Wang, which is sufficient as long as one only disperses in one direction at a time. This method has intered been described by Trueman To-).001 for treating superconvergence relations and as this is an elegent boolication I will review his work.

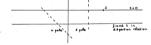
Here we write a dispersion relation at fixed t, and we can invoke the principle (P') to say

$$r_{\bullet} = \bar{H}_{\bullet} (\sqrt{-\epsilon})^{|\lambda_{\perp}^{\pm} - \epsilon \lambda_{\Gamma}^{\pm}|}$$

has no singularities in $s(\lambda_1^t = \lambda_1 - \lambda_3, \lambda_1^t = \lambda_2 - \lambda_4)$. For convenience I have multiplied by a factor

$$\int_{-t} |\lambda_i^t - \epsilon \lambda_f^t|$$

(ϵ = sign[$(m_1^2 - m_3^2)(m_2^2 - m_4^2)$]) to ensure that in the UU case f_t is analytic in t at t = 0.



So we can define residues by (5,2.1) for may poles that are present, and avaluate their residue in f_i from the TW crossing relation (2,2.8) at $s=n^2$. This is written down for large s_i and one must analytically continue (2,2.8) in a for poles which lie below the arthresholds. (Sottice we have reversed a and throughout with Chapter 2). A similar procedure works for collect in E_i .

The advantage of this application is that it is clearly

superior to invariant amplitudes, in that to write down superconvergence relations we need to know the behaviour as s → of f. In Regge theory this is very simple, being

where $\Lambda = \max(|\lambda_i^t|, |\lambda_i^t|)$.

We can then decide whether the superconvergence relations of the form \[\sigma^n \] Im ft ds converge. As usual, if (-1)^-v T = + (-1)n. Regge poles of signature T do not contribute to this relation in leading order. (v = 0 except for boson-fermion reactions in the t-channel when it is %). Also as $t \rightarrow 0$ if there is no conspiracy, the contribution of any Regge pole in leading order vanishes like $\left(-\frac{1}{2} \left\{ \left| \lambda_{i}^{t} \right| + \left| \lambda_{i}^{t} \right| - \left| \lambda_{i}^{t} \right| + \varepsilon \lambda_{i}^{t} \right\} \right)$

which will help the convergence of some relations.

(ii) Crossing

Writing
$$\overset{\text{Writing}}{\underset{\lambda_2 \lambda_1}{\forall}} = \sqrt{2J+1} \quad \underset{\underset{\lambda_2 \lambda_1}{\forall}}{\forall} \lambda_2 \lambda_1$$
(3.3.2)

the relation (3.3.1) becomes

$$H_s \sim \sum_{m} (\widetilde{Y}_{\lambda_{k}\lambda_{k}}^{T} \mathcal{D}_{\lambda_{k}m}^{T} (Q_{k}^{n})) \cdot \frac{1}{m^{k-2}} \cdot (3.3.3)$$

$$(\widetilde{Y}_{\lambda_{k}\lambda_{k}}^{T} \mathcal{D}_{\lambda_{k}\lambda_{k}}^{T} (Q_{k}\lambda_{k}))$$

where we have split up $\Gamma_{0,\theta_g,0} = R_1^{-1} R_r$, where R_1 and R_r are rotations relating the initial and final states, respectively, to some standard set of axes.

We can identify

$$\tau_{m}^{J} \frac{s_{2}}{\lambda_{2}} \frac{s_{1}}{\lambda_{1}} = d_{m}^{J} \frac{s_{1}}{\lambda_{1}} (s_{1}) \widetilde{Y} \frac{J}{\lambda_{2} \lambda_{1}}$$
 (3.5.4)

as the helicity amplitude for the reaction $1+2 \rightarrow 5$, putting $J=s_{3}$ and $n=\lambda_{5}$. In particular $\widetilde{V}_{\lambda_{\frac{1}{2}}}^{1}$ corresponds to a transition from a state with 1 and 2 along the z axis to particle 5 at rest, with $\lambda_{c}=\lambda_{1}=\lambda_{1}$.

We can now define an M function as in (2.A.7) and then cross particle 5 with 1 or 2 getting a very simple result.

To be definite, take all particles stable so that for instance $(n_1+n_2)^2>n_3^2>(n_1-n_2)^2$. Then $\widetilde{Y}_{\lambda_1}^2h$ has be continued past the threshold s = $(n_1+n_2)^2$ from sy $(n_1+n_2)^2$ where it is real (see part (iii) of this section). Let all amplitudes be continued above the top threshold, i.e.



Then on crossing 2 and 5

$$\widetilde{\chi}_{\lambda_{2}\lambda_{1}}^{J} = e^{-i\pi(s_{2}+s_{5})} \gamma_{1}^{p} \gamma_{2}^{p} \gamma_{5}^{p} \gamma_{25} \widetilde{\chi}_{\lambda_{5}\lambda_{1}}^{J}$$
 (3.3.5)

where $\gamma_{25}=\epsilon_{25}$ $\lambda_2\lambda_5^{\bullet}$ in terms of the quantities in the appendix to Chapter 2 (here λ is a crossing phase, not an helicity!).

(3.3.5) enables one to express conditions on \(\tilde{Y} \) if some of the particles are the same or if they are different, simply to relate poles in different reactions. In this latter case it is usually sufficient to know that the phase in (3.3.5) is independent of helicities.

If one of the particles is unstable, say particle 2, we are in a dilemma as (3.3.5) relates a quantity with s = $n_2^2 < (n_1 + n_2)^2$ to one with s = $n_2^2 > (n_1 + n_2)^2$, and it would inply that the former has a dynamical cut at s = $(n_1 + n_2)^2$ to correspond to the latters unitarity cut at s = $(n_1 + n_2)^2$.

(iii) Positive Definite Hamiltonian

The real analyticity of the full scattering amplitude gives a definite prediction as to the reality of the product of two \widetilde{Y}^{J_0} s. However the reality of the \widetilde{Y}^{J_0} s themselves is still undetermined up to a possible factor i. From field theory we can obtain this factor man thereby give a definite sign to the residue of poles in elastic scattering amplitudes. In order to translate this result into the helicity formalism, we temporarily piace particles of section (ii) on a Reage trajectory which we suppose to be the lesding one. Then taking a showe threshold we can make the "marrow resonance approximation" to 2-particle unitarity and assume unitarity is a survivated by this

one Regge pole. This gives at once the desired sign and shows that for my definitions $\widetilde{Y}_{\lambda_{j}, \lambda_{j}}^{-1}$ is real slowe threshold (in secondance with the approximation the small (symmanical imaginary part is neglected). But our residues occur at $s = s_{j}^{2}$ below threshold, and we say that the reality conditions can be taken below threshold in the say that the reality conditions can be taken orbital angular momentum, which are monsingular at $s = (n_{j} m_{j})^{2}$. This then says for $(n_{j} - n_{j})^{2} < n_{k}^{2} < (n_{j} - n_{j})^{2}$.

$$\tilde{\chi}_{\lambda_2\lambda_1}^J$$
 is real if $\eta_1^p \eta_2^p \eta_5^p = 1$

is pure imaginary if 1_1^p 1_2^p 1_5^p = -1

It may be proved algebraically, in case the above argument is unconvincing, that this is identical to the prediction obtained from Neinberg's Peynman rules **(4).

3.4 Form Factors

This may be regarded as an introduction to our work on Regge theory. In section 3.2 we considered the analyticity of partial wave amplitudes. The analysis of form factors is precisely the same at the thresholds $s=(n_1\pm n_j)^2$ as that given earlier where the conditions are factorizable, and so it is immediately implicable to $\widetilde{V}_{h_1 h_2}^{\lambda_1}$. Therefore, in this section, we will only consider the behaviour at s=0 and give some

examples.

(i) s = 0

As in Section 3.3(ii), we introduce the 3-particle M function. Written explicitly we have

 $M_{\gamma_1 \gamma_2 \gamma_5} = (-1)^{s_2 - \gamma_2} (-1)^{s_1 - \gamma_1} e^{\gamma_1 \sigma_1^s} e^{-\gamma_2 \sigma_2^s} \tilde{\delta}_{\gamma_2 - \gamma_1}^{\sigma}$ (3.4.1)

where $s = m_5^2$, $\gamma_1 + \gamma_2 + \gamma_5 = 0$ and M has arguments

$$\begin{array}{lll} \mathbf{p}_1 &=& \mathbf{m}_1(0,\,0,\,\sin\sigma_1^{\rm S},\,\cot\sigma_1^{\rm S}) \\ \\ \mathbf{p}_2 &=& \mathbf{m}_2(0,\,0,\,-\sin\sigma_2^{\rm S},\,\cot\sigma_2^{\rm S}) \\ \\ \mathbf{p}_5 &=& \mathbf{m}_5(0,\,0,\,0,\,1) \end{array}$$

In field theory we would assume that N obeys the principle (P). This was a satisfactory assumption before, and contained the same information whatever spinor states were used. Thus a lower-obtted spinor state is given in terms of the lower states, that we have perfously used, by:

$$lp,\dot{\epsilon}\rangle = lp,\dot{\epsilon}\rangle D_{\dot{\epsilon}'\dot{\epsilon}}(c''\frac{p.\sigma}{m})$$
 (3.4.2)

Normally the D matrix gives a polynomial in $\frac{R_{\rm eff}}{m}$ and the principle (P) (analyticity in p) is valid for all spinor states together. Unfortunately in our 3-particle cases = 0 corresponds to m_0 = 0 and a singularity in (3.4.2). Thus we will find

different predictions as to the behaviour at s = 0 of $\frac{\pi}{3} \int_{\lambda_1 \lambda_1}^{\lambda_2}$ corresponding to the infinite number of representations of the homogeneous Lorentz group in which we may place particle λ_1 . All of these results will however be consistent with the analyticity of the full smolitude calculated as

$$\langle \ell_3 \ell_4 | m | \ell_k \ell_1 \rangle = \langle \ell_3 \ell_4 | m | \ell'_3 \rangle C^{-1}_{\ell'_2} \ell_3$$

$$= \frac{1}{2} \cdot \langle \ell_5 | m | \ell_k \ell_1 \rangle$$
(5.4.3)

We will consider first the predictions of the lower M function. When $m_1\neq m_2$ we have

$$\widetilde{\mathbf{Y}}_{\lambda_2 \lambda_1}^{\mathsf{T}} \sim \sqrt{s}^{-\epsilon_1 \lambda_1}$$
 (5.4.4)

where $\epsilon_i = \text{sign } (m_1^2 - m_2^2)$.

But parity relates λ_i to - λ_i and so we end up with

$$\widetilde{\chi}_{\lambda_2 \lambda_1}^{J} \sim I_{\overline{a}}^{|\lambda_1|}$$
 (3.4.5)

The contradiction with parity (and time-reversal) of (5.4.4) could have been expected as the condition of parity conservation (2.5.11) has a $1/m_g$ in it.

In Regge theory we need only guarantee the analyticity of the amplitude to leading orders as we allow daughters to take care of the nonasymptotic behaviour. Then we find the behaviour

$$\widetilde{g}_{\lambda_{2} \lambda_{1}}^{J} \sim \frac{J_{\overline{s}} |\lambda_{1}|}{J_{\overline{s}} J}$$
 (5.4.6)

and as expected the form factor predicts a far more dramatic vanishing at $8\,=\,0$,

In the case $m_1=n_2$, we get results like (3.4.4) but with rotations by T/2 acting on the indices (see (4.2.9b)). As mentioned at the end of Section 3.2, we may introduce some new orbital angular momentum states, $\{T_i \text{ any}\}$

$$\widetilde{\mathbf{g}} \overset{\mathbf{s}'_{1}}{\mathbf{e}} \overset{\mathbf{e}}{=} \sum_{\lambda_{1} \lambda_{1}} \widetilde{\mathbf{g}}^{3}_{\lambda_{1} \lambda_{1}} \quad (-1) \overset{\mathbf{s}_{1} - \lambda_{2}}{=} e^{i \cdot w_{i_{1}} \lambda_{1}}$$

$$C \left(\overset{\mathbf{s}_{1}}{\mathbf{s}_{1}} \overset{\mathbf{s}_{1}}{\mathbf{s}} : \lambda_{1, 1} - \lambda_{2} \right) = C \left(\overset{\mathbf{g}}{\mathbf{s}} \overset{\mathbf{e}}{\mathbf{s}} : \overset{\mathbf{e}}{\mathbf{s}} : - \lambda_{1}, \lambda_{1} \right)$$

which behave like \int_{0}^{t} . Notice we have the familiar total soin s' linear combinations (Section 2.4),

I have not examined what happens when one places particle 5 in a more general Lorentz group representation, but it is of some interest to consider the case when $n_5 = 1$ and one describes particle 5 by a Lorentz vector A_f . Then, an application of the principle (P) gives, if $n_1 \neq n_2$,

$$\widetilde{Y}_{\lambda_2\lambda_1}^{\pi} \sim I_{\overline{0}}$$
 $\lambda_1 = 0$
 $\lambda_1 = \pm 1$
(5.4)

We see that the behaviour of the spin-filp and spin nonflip contributions have been exactly reversed, as compared with the prediction (3.4.5) of the lower M-function. Comparing with the Regge prediction (3.4.6), they are the same for the spin-file but the con-file has an extra vanishing in (1.4.7).

If $n_1=n_2$ a similar reversal of filly and non-flip behaviour occurs, when expressed in terms of multiuber restated by the usual equal-mass $W/2 \ v.r.t.$ $\widetilde{Y}_{k_1 k_2}^{-1}$. This has obvious relevance for the photon preferring to populate states of helicity one, among presumably one can prove a similar theorem to Misinery's on the physically acceptable Lerentz representations for the photon,

When doing fore factor calculations of $\frac{1}{2}$ exclosinge, it is consider to use an λ_{p} field for the $\frac{1}{2}$ and invoke the $\frac{1}{2}$ - photon analogy to pick out a particular from of coupling. In fact only this choice $(0_1=0_2$ at s=0 see part (ii) for notation guarantees the analyticity of the amplitude! No will comment further on this when dealing with the reaction $\pi N + \pi N^2$ in Chapter 4.

(ii) Examples

Here we calculate \widetilde{Y}^2 for some of the form factors used in Jackson and Pilkuhn⁷²), whose notation is used. (See also⁰³⁾ for some useful information).

This will be useful when comparing their coupling constants and those obtained in Regge theory where $\widetilde{\delta}_{\lambda_2\lambda_1}^J$ is parameterized.



a)
$$1^{-} \rightarrow 0^{+} 0^{-}$$

$$\widetilde{Y}_{00} = g^{-} \frac{\lambda (m_{\pi}^{2}, m_{\pi}^{2}, m_{\pi}^{2})}{n_{g}}$$
b)

 $Y_{01} = Y_{0-1} = \frac{1}{2m_c} \lambda(m_0^2, m_0^2)$

 $\widetilde{V}_{N_{1}N_{2}} = \widetilde{V}_{-N_{2}-N_{2}} = (S_{V} - S_{T}) \frac{(m_{V} - m_{A})}{m_{V}} \sqrt{(m_{V} - m_{A})^{2} - m_{V}^{2}}$ $= S_{T} = \frac{(m_{V} - m_{A})^{2} - m_{V}^{2}}{(m_{V} - m_{A})^{2} - m_{V}^{2}} \sqrt{(m_{V} - m_{A})^{2} - m_{V}^{2}}$

νε (η, -η, = δ., η, η = √Σ √(η, -η, γ. - η, ε (ζ, + ζ, γ)

r)

$$\tilde{X}_{N_{1},N_{2}} = -\tilde{X}_{N_{1}-N_{2}} = \sqrt{\frac{(m_{1}-m_{1}^{2}-m_{1}^{2})}{(m_{1}-m_{2}^{2})}} \left\{ \begin{array}{l} S_{1} & (m_{1}^{2}-m_{2}^{2}) \\ S_{2} & (m_{2}-m_{2}) & ((m_{2}-m_{1}^{2}-m_{2}^{2})) \end{array} \right\} \end{array}$$

Our crossing relation (3.3.5) enables one to extend these results to cases found by permuting any of the three particles.

CHAPTER Y

Rexxe Theory of Resonance Production

4.1 Introduction

At Introduction
In the previous chapters we have studied the inter-relation
of the vertions brumalisms that can be used in studying higher
spin particles. In this chapter we apply our techniques to the
Regge theory of resonance presention and make a direct comparison
of our calculations with experimental results. The general
methods that we have presented for dealing with higher spin can
also be applied within the framework of other theories and
problems. These include, for example, the use of the Nuncleistan
representation in the strip approximation. Avorther possible
application is given by the attempt to calculate the neutronproton mass difference by perturbing a suitchannel N over D
system of equutions in which the proton is represented as a
win and was bound state. Some aspects of the latter problem

We begin this application of Regge theory by giving some essential formalism. In Section 4.2, we work to leading order and consider compairacy briefly. In Section 4.3 we present a very liatted treatment of daughters and nonaxymptotic corrections. Then we turn to the problem of fitting experiment. In Section 4.4 we make some general comments on the experimental situation and list some technical points on the fitting procedure. After this we deal with some particular reactions: In Section 4.5, when wh $_{1236}$ and $_{\rm KN}+_{\rm KN}^{\rm N}=_{236}$ in Section 4.6, where $_{\rm N}=_{\rm KN}^{\rm N}=_{\rm$

This part of the work has been developed jointly with T. W. Rogers, who has considered the last case in more detail and also W N + $f^{0}N_{1}$, w N + $f^{0}N_{1256}^{2}$ and NN + NN_{1256}^{2} . I am grateful to C. Froggatt for discussions on Section 4.6.

The diagrams and a table of the data used are given at the end of the chapter. Here I have indicated the number of data points available and estimated the number of events so that one can excee the statistical accuracy.

4.2 Regge Theory in Leading Order

In order to consider the analyticity of the residue functions in that this chapter is anomatum transfer and \sqrt{s} is energy), we asymptotically expand a variable full amplitude. Taking the leading order, and assuming the presence of only one (or more if we have complarey) pole at $j = \kappa$, gives conditions on the residue functions. These conditions occur at t = 0 and t is also a subject to the full amplitude has the required smarticity to all orders in s. As in the spinless case the contribution of a single pole gives this correct amplyticity at the thresholds, but at t = 0 we will need

daughters in order to render the nonasymptotic terms consistent with analyticity.

In this section we will be solely concerned with the leading order contribution. There are two possible approaches. The leading order contribution of a single pole at j = < to N, is very simple and may be written

$$\lim_{t \to 0} \frac{\lambda_{2} \lambda_{1} \cdot \lambda_{3} \lambda_{1}}{2 \sin \pi \kappa} = e^{-i\pi/2} \left(\lambda_{1} - \lambda_{1} \right) \left(\lambda_{1} \cdot \lambda_{1} \cdot \lambda_{2} \cdot \lambda_{3} \cdot \lambda$$

where (a) we have specialized at once to Boson-Boson or Fermion-Fermion reactions in the t-channel.

- (b) our Regge pole has signature τ and scale factor $s_Q^{}$
- (c) $H_{\rm t}$ has been continued from the t-channel by the same route as in (2.2.8) (with s and t swopped),
- (d) partial wave amplitudes are defined by (3.2.8) for complex j and have residue $\beta_{\lambda_2\lambda_4;\lambda_3\lambda_1}$ factored as $\bigvee_{\lambda_4\lambda_2}$ $\bigvee_{\lambda_3\lambda_1}$. Finally we have put

$$\forall'_{\lambda_{j},\lambda_{i}} = \sqrt{\frac{\pi \Gamma^{1}(2\kappa+2)}{\ell^{2}(\kappa+\lambda_{i}+i)} \frac{\Gamma \log \delta}{\Gamma \Gamma_{i,j}}} \nabla_{\lambda_{j},\lambda_{i}}$$
 (4.2.2)

Asymptotically the result is in fact equally simple in terms of ${\rm H_g}^{\rm F5}$, and in order give a treatment which is independent of the masses we first consider the analyticity of

β', λ, using H_s.

(i) Direct Channel Approach (H_)

Recently it has become apparent that the interpretation of Regge theory for unequal mass scattering requires some care due to the singularity of the mapping $(s,t) \rightarrow (s, \cos \theta_*)$ as t + 0. There have been two main approaches to this problem. The first F3) assumes the amplitude has a smooth asymptotic behaviour in a pair of variables, such as (s.t) themselves. which are not afflicted by this singular manning. The second L1),G4) writes a dispersion relation in t, at fixed s, for the Regge term expressing it as an integral over its discontinuity for t > 4m2. The integrand is then asymptotically expanded in s and the integral over the leading order term done explicitly. Ignoring the possibility of fixed poles in the J plane which are suggested by the second method (these may be regarded as a special case of moving daughter poles) both methods lead to the same conclusion F6) and to the existence of subsidiary (i.e. daughter) trajectories, intercepting for t = 0 at $\kappa - 1$, $\kappa - 2$... (these will be considered in Section 4.3).

In the case of spin an appropriate generalization is to make the same assumptions about an amplitude which has no kinematic singularities or zeros in t. So a suitable candidate is \$\vec{n}_{\text{s}}\$, into which we put a smooth (in t) asymptotic behaviour. This gives in leading order for the customary route of continuation.

H, 1274 , 144 = (- 12 + 1) 3 h 3 h 3 h 1 1 1 2 2

. ϵ_{x3} ϵ_{34} ϵ_{42} λ_{x}^{*} λ_{3}^{*} $(-1)^{\frac{4}{3}} + \gamma_{x}$ $(-1)^{\frac{5}{3}} - \gamma_{4}^{*}$ γ_{1}^{p} γ_{4}^{p} τ_{1}^{p}

. 642 XLX3 (-0) ... (-1) ... 12 14 EF

Notice that the S-densed η_{L}^{c} not where the Regge pole has parity P and as in (3.2.4) η_{L}^{c} is used

$$s_{V_{3}V_{1}} = \sum_{\lambda_{1}\lambda_{2}} e^{\frac{i\pi V_{2}(\lambda_{1} + \lambda_{3})}{2}(-1)} e^{s_{3} + \lambda_{3}} Y_{\lambda_{3}\lambda_{2}}^{i}$$

 $e^{\frac{s_{3}}{\lambda_{1}V_{1}}(-X_{3}^{m})} e^{\frac{s_{3}}{\lambda_{1}V_{1}}(-X_{3}^{m})}$
(4.2.4)

and $\cos X_1^{\infty} = \frac{t + n_1^2 - n_2^2}{T_*} \qquad \cos X_3^{\infty} = \frac{-(t + n_3^2 - n_1^2)}{T_*}$

are the Trueman-Wick crossing angles with s taken to

Asymptotically the physical region boundary is t = 0 (we consider corrections to this in the next section) and so

$$g_{\gamma_3\gamma_1} g_{\gamma_4\gamma_2} \sim (\sqrt{-\epsilon})^{|\gamma_1 + \gamma_1|}$$
 (4.2.5)

and on using parity which relates $\mathbf{g}_{\gamma_4'\gamma_2'}$ and $\mathbf{g}_{-\gamma_4''\gamma_2}$ we also have

$$s_{\gamma_{3}\gamma_{1}}\ s_{\gamma_{4}\gamma_{2}}\sim\ (\sqrt{-\epsilon}\)\left|\gamma_{1}-\gamma_{1}\right|\ .$$

(4.2.6)

Y. H. is free of all kinematic singularities apart

Although the contribution of a Regge pole is asymptotically as simple in H, as in H, it will turn out to be more convenient $oldsymbol{arphi}_{oldsymbol{\lambda}_1oldsymbol{\lambda}_1}^*$ directly and so we now go on to consider this.

(ii) Crossed Channel Approach (H.)

The amplitudes g of (4.2.4) gave a very simple expression of the kinematic structure, but as discussed in subsection 3.2(1) they do not simply satisfy the nonsense conditions, which, in potential theory, are

$$\forall \lambda_{j} \lambda_{1} \sim (\kappa - J)$$
 $|\lambda_{1}| > J$ (4.2.7)

if the Regge pole chooses sense, while if it should choose

$$Y'_{\lambda_3\lambda_1} \sim \sqrt{\alpha-J}$$
 for all λ_1 . (4.2.7b)

This holds for J integral > 0, while if J is a negative integer, only the possibility (4.2.76) remains, with I replaced by

There have been many discussions lately of the status of these

results in a relativistic theory.

If the pole chooses sense, these give nontrivial conditions on the g, which are not easy to satisfy and must be imposed as extra constraints. An alternative method which we will use, is to take Y'in terms of which the rules of (4.2.7) are easily stated and impose the threshold behaviour as extra

ditions.

 $= \frac{i\pi/2(\lambda_1 + \lambda_2)}{k_{\lambda_1 \lambda_1}} \sum_{k_{\lambda_2 \lambda_1}} is real for t < 0, and more exactly we may divide out its singularities in the manner of Hara and Wang as follows.$

(a) Top Threshold

$$Y_{\lambda_{j}\lambda_{1}}^{\lambda_{j}} \sim \frac{1}{\left[e^{-(m_{1}+m_{2})^{2}} \right]^{\frac{1}{2}\left(m_{1}+m_{2}-\gamma\right)}}$$
(4.2.8a)

$$\gamma = 0$$
 if $\tau P \gamma_1^p \gamma_3^p (-1)^{n_1+n_3} = 1$
 $\gamma = 1$ if $\tau P \gamma_1^p \gamma_3^p (-1)^{n_1+n_3} = 1$

(b) Bottom Threshold m₁ ≠ m₅

$$\chi_{3\lambda_{1}}^{1} \sim \frac{1}{\left[t - (m_{1} - m_{3})^{2}\right]^{\frac{1}{2}(m_{1} + m_{3} - \gamma)}}$$
 (4.2.8b)

(c)
$$t = 0 \quad m_1 \neq m_3$$

$$8'_{\lambda_3\lambda_1} \sim (\sqrt{\epsilon})^{|\lambda_1 - \lambda_3|}$$

(4.2.8c)

(4.2.84)

(d) t = 0 m₁ = m₃

$$\chi'_{\lambda_3\lambda_1} \sim (f\epsilon)^{\gamma}$$

$$0 = 0 \text{ if } \tau P \eta^p \eta^p (-1)^{\lambda_1 + \lambda_3} = 1$$

 $\gamma = 0 \text{ if } \tau_P \gamma_1^P \gamma_2^P (-1)^{\lambda_1 + \lambda_3} = 1$ $\gamma = 1 \text{ if } \tau_P \gamma_1^P \gamma_2^P (-1)^{\lambda_1 + \lambda_3} = -1$

Only in the unequal mass case at t = 0 does (4.2.8) represent the full conditions and in the other cases we must supplement (4.2.8) with further constraints to remove the kinematic zeros. Those at the thresholds may be most easily expressed, as in the full smplitude, in terms of the perpendicular amplitudes of Kotanski [5.2.5]. This gives

$$\sum_{\lambda_1 \lambda_3} \begin{array}{c} s_3 \\ d \chi_3^{-\lambda_3}(\pi/2) \end{array} \begin{array}{c} s_1 \\ d \chi_1 \lambda_1(\pi/2) \end{array} \begin{array}{c} \chi_{\lambda_3 \lambda_1} \end{array}$$

(4.2.9a)

and as usual the negative sign is taken for the lighter particle at t = $(m_1 - m_3)^2$.

In the equal mass case at t = 0 (4.2.4) immediately gives

(4.2.9b)

(4.2.10b)

$$\sum_{\lambda_1\lambda_5} \ d^{s_5}_{\mu_3^{-}\lambda_5} (\varpi/2) \ d^{s_1}_{\mu_1\lambda_1} (\varpi/2) \ e^{i\,\varpi/2\,(\lambda_1^{\,+}\lambda_3^{\,\,})} (-1)^{\,s}\, 3^{+}\lambda_5$$

·
$$y'_{\lambda_1\lambda_1} \sim (\sqrt{-\epsilon})^{|\gamma_1 - \gamma_3|}$$

as $\chi_1^{\infty} = \chi_{\chi}^{\infty} = \pi r/2$ at t = 0 in this case.

Finally we give the conditions on Y' following from invariance under the various discrete symmetries. Time reversal has already been taken into account in using the same symbol for the counting at each vertex. Parity implies

$$Y_{\lambda_{3}\lambda_{1}}^{i} = \int_{1}^{p} \int_{2}^{p} \tau_{P} (-1)^{s_{1}^{i}+s_{3}} Y_{-\lambda_{3}^{i}-\lambda_{1}}^{i}$$
 (4.2.10a)

and if particle 1 = particle 3 $y'_{\lambda_{\overline{\lambda}}\lambda_{\overline{\lambda}}} = - \tau \quad y'_{\lambda_{\overline{\lambda}}\lambda_{\overline{\lambda}}}$

channel, (4.2.10b) has an extra phase (-1) Similarly if particle 1 = 3

$$\chi_{\lambda_{\bar{3}}\lambda_{\bar{1}}}^{*} = G(-1)^{\bar{1}} + \chi_{\lambda_{\bar{1}}\lambda_{\bar{3}}}^{*}$$
 (4.2.10c)

where the Regge pole has G parity G.

(iii) Conspiracy P1),L3)

The result (4.24), following from factorization or parity, implies that, however many Ragge poles contribute to the scattering maplitude, $H_{\rm scat} \sim \frac{|\Gamma_{\rm scat}|^2 |\Gamma_{\rm scat}|^2}{|\Gamma_{\rm scat}|^2}$. This leads one to predict a certain behaviour of the density matrix elements as t +0, which is not particularly well satisfied experimentally. So we may try to remove this theoretical prediction by allowing two or more poles to collide at the same value of κ at t = 0. There are of course many ways of obling this but a particularly interesting case is found by taking just two poles of opposite values of κ . We must easily two poles of opposite values of κ . We must easily two poles of opposite values of κ . We must easily two poles of opposite values of κ . We must easily the parity of the parity o

$$s_{\gamma_{3}\gamma_{1}}^{(1)} \ s_{\gamma_{4}\gamma_{2}}^{(1)} \ + \ s_{\gamma_{3}\gamma_{1}}^{(2)} \ s_{\gamma_{4}\gamma_{2}}^{(2)} \sim \ \sqrt{-\epsilon} \ |\gamma_{1} + \gamma_{\Gamma}|$$

which may be done by taking

$$s_{\gamma_{3}\gamma_{1}}^{(k)} \sim \frac{\int_{\overline{c}t}^{|\gamma_{1}|}}{\int_{\overline{c}t}^{-\varepsilon}} \qquad |\gamma_{1}| \neq 0$$

$$\sim \int_{\overline{c}t}^{-\varepsilon} \qquad |\gamma_{1}| = 0$$

and

$$s_{\gamma_3\gamma_1}^{(1)} = i s_{\gamma_3\gamma_1}^{(2)}$$
 for $\gamma_i > 0$

to leading order in t as t \rightarrow 0 while the coefficients of $\sqrt{-t}$ for γ_i = 0 are arbitrary.

We will apply this with Regge pole 1 as the π and g as the applitudes at the $N\bar{N}$ and πp vertices. In the first case we have $\|\nu_1\| = 1$ only but in the second $\|\nu_1\| = 0$ and 1. But in either case before conspiracy we have

$$g_{(|\gamma_i|}^{\pi} = 1) \sim f^{-t}$$
 $g_{(|\gamma_i|}^{\pi} = 0) \sim \text{const.}$

but after conspiracy .

So before compirery, the dominant (i.e. that containing the particle pole) or contribution $\mathbb{N}N=\mathbb{N}N$, $\mathbb{N}^{n}\to \mathbb{P}N$ and $\mathbb{N}N=\mathbb{P}N$ behave like $t_{i}\leftarrow t_{i}$, 1, while after compirery, $\mathbf{n}\in \mathbb{N}^{n}$ like i, $\mathbf{c}\leftarrow t_{i}$, $\mathbf{c}\in \mathbb{N}$, $\mathbf{n}\in \mathbb{N}$ the solution compirery $\mathbf{n}\in \mathbb{N}^{n}$ quantitatively, where doy't is unsitered in shape but compirery silers the expected t-variation of the density matrix slements.

4.3 Regge Theory including Nonasymptotic Corrections

In the previous section, we have chosen the t-dependence of our residues to ensure that suitable amplitudes were consistent with analyticity in leading order. We now consider whether this still holds when non-asymptotic corrections are included. Non-leading terms may be classified into three types.

- a) s^d → z^d
- b) Threshold effects

c) t = 0 and the physical region boundary. This has a part near u = 0 and another asymptotically t = 0. Both are quite important.

i) s → z (See Erratum)

As we are treating the Regge formula with some suspicion, we must consider whether to use s^{n} or z_{k}^{n} in our formulae. If we wish to avoid Regge poles with parallel trajectories κ and $\kappa-1$..., we must use z_{k} . Also z_{k} puts the cut in a sesociated with the Regge forms at z_{k} = 0, i.e.

$$2s = \sum n^2 - t - (n_1^2 - n_3^2) (n_2^2 - n_4^2) / t$$

which is more satisfactory than one at s = 0 (indicated by the asymptotic form). We have therefore used this correction in our fitting. A proper treatment of this requires the explicit subtraction of the unwanted discontinuity, and has been given for unequal mass scattering by Goldberger and Jones (44).

ii) Thresholds

As long as one uses the Arick due to Mandlestan of dropping the part of $p^{j} \ll Q^{j}$ and keep only the q^{-j-1} portion, (this is mecassary to apply Regge theory below Rej = $-\frac{1}{2}$) the nonexymptotic terms of the usual Regge form give a full amplitude with the correct amplitudity, as long as the residue functions have the form determined swithoutically. This statement is

only exactly true at the lower threshold, as unitarity is well known to require a condensation of poles at $\hat{L}=-1$ at the top threshold. This appears rather high in the j plane $(i.e.\ j=\max\{s_1+s_3,\ s_2+s_4\}-1)$ for spinning particles but we will ignore it.

The thresholds become rather more interesting when in part iv) of this section we add daughter trajectories with

- iii) t = 0 and the Physical Region Boundary: Direct Channel Method
- An asymptotic behaviour which is consistent with analyticity, for fixed s, and all t is obtained by placing

$$H_{g}^{P_{3}P_{4}:\,P_{2}P_{1}} \ = \ B_{g}^{P_{1}\,P_{1}} \left\{ \left. \frac{1}{B_{g}^{P_{1}\,P_{1}}} \right|_{\text{asymptotic form}} \right.$$

. equation (4.2.3) + any corrections of $O(s^{4.1})$ (4.3.1)

residues determined at t = 0.

This is unsatisfactory for two reasons:

(a) It gives Regge polefwith J differing by integral values for pilt which would not be sensible in some potential theory limit (and relativistically summing ladder graphs does not give parallel trajectories ⁵³), ⁵⁴⁵).

(b) It does not give the correct analyticity in s. This is

probably not usually important, but becomes so if the Regge pole creates a particle at t = n^2 . Then the formula (4,3,1) by no means ensuires that at t = n^2 the residue in n^2 has a sependence $< d^4(q_2)$. This enhancement of the nonmalyticity in a becomes particularly apparent for the \forall Regge pole and I have verified on the computer that, whereas is x = 5 GeV the natural non-mamphotic corrections are quite small (see Section 4.3), formula (4,3,1) gives results differing by 50% from the avenuatority value.

A nore sensible method, which answers the first objection, may be found in the Q_t symmetry approach developed by Freedom and Hangers. This model is only applicable netwestly to equal mass scattering but may have a more general use as in a Wick related sether-abjecter equation the masses become equal in pairs at + 0. It seems more complicated to apply than a simple Regge model, other than just at t = 0, as it only applicable and without symmetry-breaking to the coefficient of $\int_{-T_t}^{T_t} Y_t^{r-1} Y_$

The second objection may be overcome by using invariant amplitudes as these remove the singularities in s and t simultaneously. However we saw in Chapter 2 that they contained

no more information on the analyticity at t=0 than the principle (P') and so there is no need to use them in a theoretical treatment of daughters and the behaviour at t=0. As no general simple formula is known for them they are not even useful as a convenient observable device.

iv) t = 0 and the Physical Region Boundary: Crossed Channel
Method

The method of the previous subsection generated results which were non-singular at t = 0. However as we discussed there, they had certain disadvantages and so we will now study the same problem keeping rather closer to the natural Regge form, for all mass values, \vec{l}_{ij} is a mutable emplitude to study analyticity at t = 0 and on the physical region boundary. However in the case of UU scattering we can use (ef. Section 3.3)

$$r_{t} = R_{t}^{\lambda_{2}\lambda_{4}t\lambda_{3}\lambda_{1}}(\sqrt{-t})^{|\lambda_{1} - \epsilon \lambda_{f}|}$$
(4.3.2)

where

$$\epsilon = sign \left[(m_1^2 - m_3^2) (m_2^2 - m_4^2) \right]$$

which is non-singular at these two points and more convenient for Regge theory.

Me will say something later about US and SS scattering. The full contribution of a Regge pole to $f_{\rm t}$ is

$$f_t = f_t \Big|_{asymptotic} \cdot \left[\frac{T_{13}T_{24}}{st}\right]^{\kappa-\Lambda} \Big[\frac{1}{2}z - \frac{1}{2}\right]^{\kappa-\Lambda}$$

(4.3.3)

third term and

where the second term is just 1/the asymptotic value of the

Because $z \to 6$ as $t \to 0$ for all s, z - 6 is a satisfactory quantity to appear and $(\frac{z-6}{6})$ and t are a pair of variables non-singularly related to s and t as $t \to 0$.

Working to O(s -1) the hypergeometric function becomes

$$1 + \frac{(\Lambda - \kappa)(\lambda - \kappa)}{-\kappa t} \cdot \frac{t}{(1 - \epsilon z)}$$
 (4.5.4)

which has a j/k pole as $t \to 0$. This must be cancelled by one or nore daughter trajectories. In order to make the study nontrivial we will suppose there is but one daughter trajectory (at $j = \kappa - 1$) which has factorizable residues. It therefore must contribute to f_k a term which is proportional to the 1/k part of the 2/k1 and hence adds to (4.).4) a term

where from factorization we have $|\lambda|$ and not λ . So the joint contribution of parent and daughter at t = 0 may be obtained by replacing (4.3.4) by

$$1 + \frac{(\Lambda - \kappa)(\lambda - |\lambda|)}{-\kappa t} \cdot \frac{t}{1 - \epsilon z}$$
 (4.3.5)

 f_t must be non-singular and there is still a 1/t in (4.3.5) if λ negative. However factorization of the parent trajectory implies from (4.2.6) that as t \rightarrow 0,

$$f_t$$
 asymptotic \sim const if λ positive \sim $t^{|\lambda|}$ if λ negative

This behaviour when multiplied into (4.3.5) gives a non-singular form to $f_{\underline{e}^{\star}}$

In terms of χ' the parameters of the daughter are given by:

$$\alpha_{(q)}^{(q)} = \alpha_{(p)}^{(q)} = \alpha_{(p)}^{(p)}$$
(to be $\alpha_{(q)}^{(q)} = \alpha_{(p)}^{(q)} = \alpha_{(p)}^{(q)} = \alpha_{(p)}^{(q)}$

teo be concinately

. - - -

(4.3.6 continued)

$$\chi_{\lambda_{3}\lambda_{1}}^{\tau(d)} = \chi_{\lambda_{3}\lambda_{1}}^{\tau(p)} \frac{(|\lambda_{1}| - \kappa)}{\sqrt{-2t \kappa s_{0}}}, \tau_{13}, \epsilon_{1} \text{ (as } t \to 0)$$

where $\epsilon_1 = \text{sign } (n_1^2 - m_3^2)$ (and we have given the parent and daughter the same scale factor $s_{\alpha\beta}$).

Towerk to order s $\stackrel{<}{\sim}$ 2 and lower is only well defined for the most singular part in t, as the other singular terms are affected by the difference between the daughter and parent parameters may from t = 0. Thus (4.7,6) cannot hold for all t as it controllets the threshold behaviour at $T_{1,2} = 0$. (This cours even is the spalless case when Y is non-insquier at thresholds). However one can easily give a formal extension of the analysis to all orders in a by using the identity. (Marriers and Gunnard), equation (5.7,1),

$$1 = \sum_{r=0}^{\infty} \frac{1}{r!} \frac{\left(\Lambda - \kappa\right)_{\Gamma} \left(\lambda - \kappa\right)_{\Gamma}}{\left(r - 1 - 2\kappa\right)_{\Gamma}} \left(-x\right)^{\Gamma} {}_{2}F_{1} \left(\Lambda - \kappa + r, \lambda - \kappa + r, \kappa\right)$$

which is simply a complex form of the C.G. series for

$$\mathsf{d}_{\mathsf{A}-\mathsf{J}_{\mathsf{b}},\mathsf{J}_{\mathsf{a}}}^{\mathsf{J}_{\mathsf{a}}}(\boldsymbol{e}_{\mathsf{t}}) \quad \mathsf{d}_{\mathsf{J}_{\mathsf{b}},\;\lambda-\mathsf{J}_{\mathsf{a}}}^{\mathsf{J}_{\mathsf{b}}}\left(\boldsymbol{e}_{\mathsf{t}}\right) \quad \left(\mathsf{J}_{\mathsf{a}}+\mathsf{J}_{\mathsf{b}}=\infty\right)$$

Putting $x=2/(1-\varepsilon z)$ and then multiplying through by $(-1/x\varepsilon)^{\kappa-\Lambda}$ we get the sum of parent pole and daughters on the

right hand side equal to the, well-behaved as $t \to 0$, function $(kx - kt)^{\frac{1}{n-1}}$. This, as for the first daughter, only works exactly for the positive λ but, as before, the asymptotic vanishing of f_t ensures that the daughter residues determined for $\lambda > 0$ are sufficient to ensure smalltitity for $\lambda < 0$.

We had to use the above device rather than expand (4.3.3) further as the residue of the second daughter is determined by terms from both the parent and the first daughter.

So the residues of the daughter at $\ensuremath{\omega^{(p)}}(0)$ - r are given by

$$Y_{\lambda_{3}}^{(d)} = Y_{\lambda_{3}}^{(p)} = \frac{(|\lambda_{1}| - \epsilon)_{\Gamma}}{\sqrt{\Gamma(\Gamma - 1 - 2\epsilon)_{\Gamma}}} = \left[\frac{\tau_{13} \epsilon_{1}}{\sqrt{\tau \epsilon_{0}}}\right]^{\Gamma}$$
 (4.5.7)

Again we have one and only one daughter trajectory at a - r (r integral) and if (4.3.7) held as identity in t (and the daughter trajectories were parallel to the parent away from t = 0) this would ensure that f_c had no singular terms. However as discussed above this is impossible and the above analysis only holds for the most insurals terms in t.

We will now consider UE and EE scattering. UE scattering seems the most difficult but a treatment to order $s^{4-\ 1}$ may be given as follows.

We now state our analyticity in terms of the Trueman-Wick crossing relation and the non-singularity of \vec{R}_g . As we are only interested in t \approx 0 and the physical region boundary and

not the other t thresholds we may introduce the total spin s', which as mentioned before (Section 2.4) always diagonalizes the behaviour we want here, and simplify the statement of analyticity by outting

$$m_1 = m_2$$
: $H_0^{\sum_i \lambda_{ij} - \lambda_{ij}} = \sum_{\lambda_i = \lambda_{ij} = \lambda_{ij}} C(s, s_2 s_1^i : \lambda_{ij} - \lambda_2)$
 $(-1)^{\sum_i \lambda_{ij} - \lambda_{ij}} = H_0^{\lambda_i \lambda_{ij} + \lambda_{ij} + \lambda_{ij}}$

which satisfies

(4.5.8)

$$\sum_{\lambda_{\underline{i}}} H_{\underline{t}}^{\lambda_{\underline{2}} \lambda_{\underline{i}}; \lambda_{\underline{i}}} d_{\lambda_{\underline{i}} P_{\underline{i}}}^{s_{\underline{i}}} (X)$$

vertex, the perpendicular amplitudes

has only the physical region boundary behaviour $\prod_{r \in cos 0} \sqrt[r] h_r^{r_r} f_r \mu_1^r$ where X may be taken as X_1 or X_2 (or $X_1 = \frac{X_2}{L}$, for that matter). Because in US scattering, the physical region boundary differs from t = 0 by terms proportions to $1/\delta^2$ ($h_r + \epsilon_r \mu_1$) for the purpose of our malysis to $0(\delta^{d-1})$.

From (4.2.9b) it is useful to introduce, at the equal mass

$$P_{e_{1}^{1}}^{V_{1}} = \sum_{\lambda_{1}\lambda_{3}} c (s_{1}s_{3}s_{1}^{1} : \lambda_{1}, -\lambda_{3}) (-1)^{s_{3}-\lambda_{3}} e^{-i\pi/2} \lambda_{1}$$

$$(4.5.9)$$

 $y'_{\lambda_3\lambda_1}$ $d_{\lambda_1\gamma_1}^{s_1'}$ $(\pi/2)$

(4.3.11)

Instead of z-4 which was an appropriate variable in the UU case, it is now sensible to expand in terms of z. Then if we assume that there is again only one desights whose residue at the unequal mass vertex is given from factorization by (4.5.6), we can show that the daughter residues at the equal mass vertex are given by

$$P_{k_{i}}^{k_{i}(k)} = \frac{-m}{\sqrt{k + k_{i}}} \left(\frac{(k_{i}^{k} - |y_{i}| + 1)^{k_{i}}}{(k_{i}^{k} + |y_{i}|)^{k_{i}}} \frac{(y_{i}y_{i})}{(y_{i} + y_{i})^{k_{i}}} \right) \frac{(y_{i}y_{i})}{(y_{i} + y_{i})^{k_{i}}} \left(\frac{y_{i}y_{i}}{y_{i}} \right) \frac{(y_{i}y_{i})}{(y_{i} + y_{i})^{k_{i}}} \frac{(y_{i}y_{i})}{(y_{i} + y_{i})} \frac{(y_{i}y_{i}$$

while for $\gamma_1=0$ the daughter can have any residue as long as it vanishes like \sqrt{t} to cancel the $1/\sqrt{t}$ in $\gamma_1^{*,\{d\}}$.

It may be verified that the equal mass conditions ((5.2.11)) are also satisfied to $O(s^{4-1})$ by the same daughter with factorized residues given by (4.5.10).

As an example we can take the well known case of nucleonnucleon scattering. In terms of the perpendicular and s' notation, which is rather cumbersome in this special case, we have three types of complings:

$TP = + : G(-1)^{I}P = +$

s' = 0 one coupling which is irrelevant here
s' = 1 one coupling p⁰ = 0

(4.3.11)

$$p^{0} = i \sqrt{2} \quad \forall_{1}$$
 $p^{1} = p^{-1} = 0$

 $\begin{array}{cccc} T & TP = - & : & G(-1)^T & P = - \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & & \\ & & \\ & & & \\ & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & & \\ & &$

$$p^{0} = 0$$
 $p^{1} = -p^{-1} = \chi_{0} / \sqrt{2}$

Here χ'_{λ_i} stands for

$$\sum_{\lambda_1 + \lambda_2 = \lambda_1} C_i(s, s_2, s_i' + \lambda_{i_2} - \lambda_2) = (-i)^{s_2 - \lambda_2} = \chi'_{\lambda_2, \lambda_i}$$

for $s_{\underline{1}}^{\,\prime}$ = 1 and so are directly the residues to be inserted in the standard relation (3.2.10).

Our rules (4,3,10) show that only one type has a daughter to j=k-1 with singular residues. This is type II (sasociated with the A_j or p seems) with a type III (" n, γ) daughter. This is a particular case of many solutions known to (5,1,2,10). One sassuption that there is only one daughter covering all reactions has picked out a unique solution.

This method of tackling the UE and EE cases seems incopole of easy extension to $O(s^{-2})$ and lower. Also it makes it uppear rather alreadous that the daughter has factorizable residues, while in fact this should have been expected. Thus the only reason UE and EE reactions are different from the general UU case is that the thresholds $(n_1-n_2)^2$, $(n_2-n_1)^2$ move down and coalesce with t = 0. Now the threshold conditions

in the UU case required no extra daughters and so one would not expect extra once to be produced when t=0 happens to coincide with $t=(a_1-a_2)^2$. This can be made the basis of a quantitative study of EE and UE rections as follows. Thus take the UU expressions for the daughter residues and write them in terms of amplitudes which have no kinematic singularities or zeros at the thresholds, as emphasized before, even in the spiniese case, this is not unconstit from the natural factors of (4.5, 1.6) and (4.5, 7.7). Then letting $n_1 + n_2$ we would achieve the correct results for UE and Excattering. This is straightforward for the most singular term in t and although one does not know an explicit general solution it is not very difficult in practice for n_2 when we can be early one n_3 or n_4 where n_4 is not required.

The above analysis may seen rather complicated but in practice it would probably be sufficient to include mon-saymetotic terms only if they were singular compared with the parent, and destroyed the asymptotic vanishing of f_s. In our model which makes the arbitrary assumption of one doughter at each « - r, this destruction occurs in a controlled way. Meanly a term supportional to ghill loses one factor of t each time you go down one power in a until you reach a black of the supportional to ghill lose one factor of t each time you go down one power in a until you reach a black of the supportion one can be a final that the further non-saymptotic terms are not more singular. In analysing the experimental situation one can bear in mind that non-saymptotic terms may be more singular than the above model suggests.

Finally we note that the simple prescription of just dividing out the physical region boundary behaviour

(4.3.12)

although consistent with smalylicity at t = 0 would lead to non-asymptotic contradictions with analylicity at the thresholds. As experiments are consisted searer the physical region boundary than the thresholds this may not be important. The natural leages non-asymptotic terms restore consistency at thresholds and in fact make the corrections may from the physical region boundary smaller. Thus (4,3,12) gives corrections of Oil/s) relative to leading order in physical quantities but as

$$d = \frac{M}{\lambda_{\perp} \lambda_{\Gamma}}(z) = Asymptotic form \left\{ 1 + \frac{\lambda_{\perp} \lambda_{\Gamma}}{MZ} + \dots \right\}$$

(where asymptotic means asymptotic in x not s, here see subsection (i)) the non-asymptotic correction belowes like a pole of opposite \tau for the leading order and so does not interfere with it in calculating physical quantities. So the only 1/s correction is the daughter-leading order interference.

4.4 Experimental Fitting

(i) The quantities which we must try to fit are:

$$\frac{d\sigma}{dt} = \frac{0.3893}{16 \text{ m s}_{12}^2} \sup_{\substack{\text{sun over final and average over initial}}} \left| \prod_{k=1}^{N_2} \lambda_k; \lambda_j \lambda_k \right|^2 \qquad (4.4.1)$$

and the density matrix elements. These are usually referred to a set of axes popularized by Jackson, and

$$\rho_{mn'} = \frac{\hat{\rho}_{mn'}}{\text{Tr } \hat{\rho}_{mn'}} \qquad (4.4.2)$$

here

$$\hat{\beta}_{mn}, = \sum_{\lambda_1 \lambda_2 \lambda_c} (-1)^{m-m} H_c^{\lambda_2 \lambda_c; m \lambda_1} H_c^{\lambda_2 \lambda_c; m' \lambda_1} *$$

for particle 3 decaying while for particle 4 the formula is the same with 3 ** 4 but without the factor (-1)****. We must be careful about the definition of our y axis. I have taken it along \$\rho_{0}^{1.00}\$ x \$\rho_{0}^{0.00}\$ x fo both particles 3 and 4. Everybody follows this procedure for particle 3 but for particle 4 (which will be the \$\rho_{136}^{0.00}\$) there appears to be some confusion. The KN G.E.R.N. experience and Jackson and Pilluhund²³ make the opposite choice but I have assumed that the other experiences have made my choice. The sign of Re\$\rho_{3}^{0.1}\$ is altered on changing the direction of the y axis and as this is small, little qualitative difference will be made to the first if my assumption is not

correct.

As g_{min} is given by t channel amplitudes it is strongly affected by their kinesatic singularities (especially for the \mathbf{N}^2) which may obscure the dynamical information. This may be an argument against using this particular set of axes. Thus we saw carlier in Section 4.3 that the simplicity following from a single Regge pole exchange was as easily stated in terms of \mathbf{N}_{ij} as \mathbf{n}_{ij} .

(ii) The reactions we will be interested in and some of the Regge poles which may be expected to dominate are given in the table below.

Resction	τP = +	τρ = -	
πN → πN* KN → KN* πN → qN*	g (g') g ('g') h ₂	not sllowed $ \begin{matrix} A_{\underline{1}}, \pi \\ B \\ A_{\underline{1}}, B, \pi \end{matrix} $	
π N → p N* π N → ω N* κN → κ*N*	Λ ₂ β g , Λ ₂		
	Type. I	Type II	Type III
πn → g n kn → k*n	Α ₂ , ω Α ₂ , p, P, P', ω	A ₁	म В,π

The division into the two values of TP is useful as from (4.2.10s) we see that in leadin; order poles of opposite TP do not interfere in either $d\theta'$ 0t or density matrix clements. Also when there is an NR vertex the three types of (4.5.11) do not interfere.

We notice that, except for the W, the poles of high intercepts -60 have CP = -8. Concrimentally this is born out in our reactions except at the lowest energies where we definitely need more CP = -8 then the W provides. This could either be due to a failure of Reage pole theory (e.g. the background integral is important) or that the A_1D and B mesons with probably somewhat lower intercepts are becoming important. I will usually consider all data above $p_{120} = 2$ GeV, which is an unfortunately low limit but the most reliable experiments lie in the 2 to B GeV region. The errors on the density matrix elements are especially large at high energies and fits involving only the high energy data are largely determined by dP/di. At lower energies where such more encurate density matrix girls giarge contributions to X^2 as they are very sensitive to the exchanged X.

In the non-charge exchange reactions such as $\pi^\pm p + g^\pm p$ there is the well known forward scanning bias, which means that $d\sigma/dt$ and the density matrix elements are unreliable for small t. I have deleted all such suspicious points although it would appear that some experiments have made the necessary

corrections.

(iii) Characteristics of the Exchanged Poles

The density matrix elements for "N $\Rightarrow N$ and ES $\Rightarrow K^2N$ are especially sensitive to the value of \uparrow of the exchanged pole as will be described in more detail in Section 6.6. For h^2 production the density matrix elements are not so critically dependent on T. (See Section 6.7). Also the density matrix elements will enable one to detect if a pole chooses sense or nonecome when it goes through zero mert \uparrow = 0.

However there is one general comment that may be made about the shape of de/dt produced by these poles. Thus W has a nearby pole; and • a reasonable slope ≪ . quite a near pole and a strong spin-flip vanishing at t = - .6. So these two noies produce do/dt's which fall off rapidly with t. W may be similar but the evidence is less conclusive. These conclusions will be somewhat modified for ø and w in KN → K*N and πN → oN where we have only a spin-flip coupling at the meson vertex and the associated t = O vanishing can give a broader dr/dt. A, and P' however produce a characteristic broad cross section, as the lack of dips in their reactions may be because they choose nonsense at $\infty = 0$ so that the vanishing (4.2.7) cancels the ghost and there is no residual spin-flip vanishing. (iv) I have written programs which fit the experimental do/dt's and density matrix elements to the leading order formulae for on arbitrary number of noise in the reactions of interest here,

The residue functions y'_{1-1} were parameterized as poly-

nomials with an overall exponential dependence in t being given by a choice of so. X² was minimized with respect to the coefficients of the polynomials by evaluating the derivatives in the usual way. The threshold constraints (4.7-30) gave linear relations among these coefficients which were conveniently incorporated using Lagrange multipliers. The parameters co., 2, and s., for each pole were fixed during each linear coefficients which were conveniently incorporated using Lagrange multipliers.

In the next three sections we will describe the results of these fits for the three classes of reactions. We can summarize the conclusions by againg them is in general reasonable agreement for any given reaction. The strong correlation in the determination of parameters from any one reaction renders tests of factorization difficult. However there appear to be some significant differences in the parameters necessary in different reactions (especially for the "W") which may indicate the presence of cuts in the j plane.

Lasty I should say that there have been several Megge pole analyses⁷⁵) of the reactions of interest. I can and no more to this work on the energy variation of "cotal (see Marrison⁹²) and df/dt. However the previous work seems to contain technical errors especially in the treatment of density matrix elements.

4.5 πN →πN₁₂₁₆ and KN → KN₁₂₁₆

(i) <u>General</u>
There are four couplings at the $8n^4$ vertex into which we may coalesce the single $\pi\pi$ coupling. At $t = (m_{11} - n_{12})^2 \approx .09$ Gev, there are from (4.2, 9a) four constraints while we will impose the two further constraints at $t = (n_{11} - n_{12})^2$. The closeness of the threshold and the number of the conditions would seem to make them quite important. In order to illustrate the constraints, suppose that the helicity double-flip term $(1 - n_{12})^2 \approx 0$. Then the remaining residues \sim constants at $(m_{11} - m_{12})^2$ or Theorem 7.1/ $(n_{12} - n_{12})^2$ and satisfy the one relation

$$\sqrt{3} \left(\chi_{\frac{1}{2}\frac{1}{2}}^{1} - \chi_{-\frac{1}{2}\frac{1}{2}}^{1} \right) = \chi_{\frac{3}{2}\frac{1}{2}}^{1}$$
 (4.5.1)

This clearly indicates that although the helicity-flip terms vanish at t = 0 they are related in magnitude to the non-flip term at t = .09. Thus we should not be surprised to find evidence for strong spin flip terms in the density matrix elements.

There are three density matrix elements measured experimentally: p_{53} , Re p_{5-1} and Re p_{51} (where the index is 2m not m).

If we take $Y_{\frac{1}{2}\frac{1}{2}}^{1}$ = 0 and assume the other residues are sufficiently constant to be able to apply (4.5.1) away from

t = .09 we find

$$p_{33} = 3/8$$
 $p_{3-1} = \sqrt{3}/8$ $p_{31} = 0$ (4.5.2)

which is the well known Stodolsky-Sakurai distribution which we see is to some extent determined solely by kinematic effects. In particular this distribution and a turn over in de/dt as $t \rightarrow 0$ (i.e. $\chi_{3k}^{+} = 0$) are kinematically resited.

As pointed out by Bialas and Kotanski if only one pole is exchanged, we have

$$g_{33}$$
 $g_{11} = (\text{Re } g_{31})^2 + (\text{Re } g_{3-1})^2$ (4.5.5)

This is well satisfied in both reactions although there is insufficient published information to evaluate the error properly.

Finally we note that non-asymptotic corrections are expected to be quite small because one vertex is spinless and the daughters have residues which are no more singular than the leading term. Also, even at plan = 1.59, the physical region boundary differs from t = 0 by only .001 Gev, (for $\pi N \rightarrow \pi N^*$). (11) m+p + m0N+++

Of the data used only that at 4 Gev has reasonable statistics. In this reaction we only have one important pole exchanged: the p pole with & a .58 and & . . The density matrix elements show an increasing and large amount of spin flip terms

as t moves away from the physical region boundary, which we expect from the arguments in (i). Thus one might expect the Schoolsky-Sakuria assumption $Y_{1,2}^{i} \equiv 0$ to be unnecessary and that the situation could be described by simple constancy assumptions subject to (4.5.1) (or more correctly, the full in the conditions with $Y_{1/2,2} \not\equiv 0$). Shewere the data move a dip in dryfit in the forward direction which is extremely difficult to fit due to the shary fall off of the g exchange contribution. Thus all my fits needed $Y_{1/2}^{ij}$ to be much smaller than kinematically expected and gave theoretically a fractionally larger $p_{1/2}^{ij}$ than the experiments value for low t.

The strong spin flip terms lead one to predict a dramatic dip in the cross section at tw. 6 where the $\mathfrak g$ trajectory passes through zero. Unfortunately the best dip in the data occurs at tw. -15. We can improve the agreement here by invoking the $\mathfrak g$ trajectory, which is airsedy needed in one model of TN C harge exchange, to fill in the predicted dip. The $\mathfrak g$ maintens were insensitive to the introduction of the $\mathfrak g^*$ as long as you kept the interesting $\mathfrak G$.2 so that the energy dependence of the data determined the $\mathfrak g$ as the dominant contribution.

The 4 and 8 Gev data had 37 data points and with 3 degrees of freedom for the $\mathfrak g$ and 1 for the $\mathfrak g^*$ we found a χ^2 . between 40 and 50 (depending on $\mathfrak u_0$ and the $\mathfrak g^*$ parameters). Figure 4.1 shows a typical fit to the 4 Gev data,

From the energy variation of σ_{tot} the lower energy data appears of suspect normalization. Allowing de/dt scale factors determined at 3.54 Gev to be = 2 and at 2.75 Gev to be = 1.5 we had a χ^2 from 80 \rightarrow 100 on a grand fit, to all 80 data points, using the same number of theoretical parameters as above. This is a measure of the poor statistics of the lower energy data rather than the applicability of the theory at low energies.

The important nature of the four kinematic constraints might lead one to expect a dependence on the mass of the N* used, On altering m_{N+} to 1.15 Gev (so that $(m_N - m_{N+})^2 \implies .05$) we found no change in χ^2 as there were no density matrix elements for small enough t for this shift to be significant. However it may be significant that the kinematically expected value of X't, goes down, after this mass change. (111) K*p + KON*++

In this reaction p and A, are expected to be the dominant poles but from SU, and the WN + WN cross section the A, is expected to be dominant. So we can first try a fit with just A., and as discussed in Section 4.4 we have an immediate qualitative explanation of why KN -> KN* has a much broader der/dt than πN → πN*. The experimentalists have already done an analysis of the energy variation of this and found & . = .4 ± .22, which is excellent, but obtain also the unfortunate result that $\alpha_0^{\dagger} = 1.72 \pm .33$. So we found for the A₂ parameters $\mathbf{x}_0 = .35$ $\mathbf{x}_0' = .36$ rather bad agreement but for $\mathbf{x}_0' = .6$ (the density matrix elements imply that λ_0 must choose noneans at $\mathbf{x} = 0$ for this large slope) there is a great improvement, although the $\mathbf{t} = -3$ $p_{1,0} = 5$ Gev data point was still too large theoretically (by three standard deviations). Thus on 50 data points and 3 degrees of freedom for the λ_2 we found a $\chi^2 = 310$ in the first case and $\chi^2 = 90$ in the second.

We then added g archange with parameters determined from $\forall n \to n^2$ analysis and with its SU, ratio. This leads to roughly equal g and λ_1 contributions near t = 0 but with the rapid g, fall off and broad λ_2 the λ_3 dominates away from the 0. This reduces X^2 by about 10 but it did not start the systematic disagreement discussed above. Increasing the contribution led to \hat{r} rise in X^2 due to the theory being too large for small i. So one may say the data is consistent with a g contribution ranging from nothing up to 1.5 times the SU prediction.

If the rapid fall off with to f the § Gov experiment is confirmed at higher energies it would suggest that the amount of $\mathfrak g$ exchange has been underestimated in the above analysis. In this respect we may note that the similar reaction $\pi \pi \to \eta \pi^*$ which only has Λ_2 exchange shows the characteristic broad cross section in the data $^{10.2}$ 0 at 8 Gov.

The data was consistent with A_2 parameters rather similar to the g (i.e. small $Y_{\frac{1}{2}\frac{1}{2}}$, $Y_{-\frac{1}{3}/2}$, in agreement with the

Stodolsky-Sakurai prediction) but reasonable fits can be obtained without this feature. Whereas the $\mathfrak g$ can have a sensible $\mathfrak s_0$ — 1 the $\mathfrak d_2$ needs a small value $\mathfrak s_0$ — 1.1 in order to pull its $d\gamma d$ to find for large t. The exact value of $\mathfrak s_0$ depends on $\mathfrak u_0^{\lambda}$ and whether the A_{λ} chooses sense or nonsense when it goes through zero (either was consistent with the density matrix elements for the small slope $\mathfrak u_0^{\lambda}$ e. 3λ).

Figure 4.2 gives the results of one of the fits, using ${\rm A}_2$ only, for the 2.97 GeV experiment.

4.6 TN → pN and KN → K"N

(i) General

This class of reactions have the following features:

(a) Three charge states of which the charge exchange reaction

- picks out I = 1 exchange. In both examples the two non-charge exchange reactions, e.g. $\pi^{\pm}p \to g^{\pm}p$, have cross-sections which are compatible with equality.
- (b) There is a large amount of data although not over a wide range of energies.
- (c) The coupling structure at the two vertices is summarized below.

	-(1)T-	Туре	Nonconspiracy		Conspiracy	
"	0(-1)		Meson Vertex	NÑ Vertex	Meson Vertex	NÑ Vertex
٠	+	1	¥1=-¥1- √€	¥'++=¥'_~const	¥'i ∼const	¥;, ~Æ
			¥1 = 0	Y'_=Y'_ ~ 1€	¥ 0 = 0	¥'_~const
-	+	11		A,++=A;-=0		
			¥1=¥1-1~ √€	¥'_=-¥'_~const		
-	-	111	¥ o ∼ const	¥;+=-¥;_~√€	¥'₁ ~ const	¥¦~const
				χ ⁺⁻ =χ ⁻⁺ =0	80~5€	₹. <u>-</u> o

The structure of two couplings at the NR vertex for a type I pole can be ignored initially as we have no polarization data and always sum over the NR indices.

(d) There are again three density matrix elements measured experimentally: \$\oldsymbol{p}_{00}\$, \$\text{Re} \oldsymbol{p}_{1-1}\$, \$\text{Re} \oldsymbol{p}_{10}\$.

From (4.4.2) we find the usual results

$$\tau P = + : g_{10} = g_{00} = 0$$
 $g_{11} = g_{1-1} > 0$
 $\tau P = - : g_{00} \propto g_{12}^{*2} > 0$ $g_{11} = -g_{1-1} \propto c g_{12}^{*2} > 0$ (4.6.1)

This implies that if we form \mathfrak{g}_{00} , \mathfrak{g}_{11} , \mathfrak{g}_{1-1} and \mathfrak{g}_{11} , \mathfrak{g}_{1-1} we pick out the relative (leading order) contributions of the $\mathfrak{TP}=-, Y_0^{-2}$, $\mathfrak{TP}=-, Y_1^{-2}$ and $\mathfrak{TP}=+, Y_1^{-2}$ couplings respectively. The higher energy data have too large errors to make this very useful at the present although plotting these out does not contradict one's expectation of their energy dependence (e.g. $\mathfrak{TP}=-$ having a higher intercept than $\mathfrak{TP}=-$).

The above holds however many mesons are exchanged but if only one tP = - Regge pole is present we have analogously to (4.5.3)

$$\mathbf{p}_{00}$$
 ($\mathbf{p}_{11} - \mathbf{p}_{1-1}$) = $2 \mathbf{p}_{01}^2$ (4.6.2)

The experiments satisfy this to within their internal inconsistencies. (The 2.72 Gev g^0 n data badly violates it but the 2.7 Gev data agrees well).

Now let us consider the expected t behaviour of the density matrix elements. As usual we take t = 0 and the thresholds separately and begin with t = 0. As we have an UE case H_L is singular at t = 0 but we can overcome this as the equal mass indices are summed over. Thus we can introduce the Truemanwick crossing matrices into (x.12), so that b²_{ma} is given by the same expression but with the NR indices belonging to the s not the t channel. (We also used this partial rotation in (4.3.8)). The hybrid amplitudes now separating in the formula for \hat{p}_{min} have the advantage of only having the physical region boundary singularity in t. So we find the following behaviour of the density matrix elements at t = 0 and the physical region boundary.

	Exact	Asymptotic———					
		General Result	Noncon- spiracy Regge Result	₩ noncon- spiracy	conspiring W	conspirator	
ŝ ₀₀	const	const	const	t/(t-m _f ²) ²	t/(t-n ² / _e) ²	. 0	
ĝ 01	sin ^k e	t ^{kj}	t ^{3/2}	t ^{3/2} /(t-m ₊ ²)	t /(t-m2)	0 .	
ĝ ₁₋₁	sin 0	t	t	t	const	= - T contribution at t = 0	
ĝ.,	const	const	·	٠	const	= + T contribution at t = 0	

I have appended the nonconspiring and conspiring behaviour for the π and put in its pole as this occurs at essentially t = 0. The observed density matrix elements are of course

p = \$/Tr\$.

We notice that the behaviour in the presence of conspiracy is the most singular allowed while in the nonconspiratorial case we have an extra vanishing in \hat{p}_{01} and \hat{p}_{11} , and would expect \hat{p}_{01} to be small and $\hat{p}_{00} = 1 - 2 \hat{p}_{11}$ to be large

Now consider the thresholds. There are no conditions at the meson vertex for the tP = + Regge poles to satisfy. Let us parameterize the W couplings at the meson vertex by

$$Y_0' = \frac{\Lambda}{T_{1,5}}$$

$$Y_1' = \frac{B \kappa_W / E}{T_{1,5}}$$
(4.6.3)

We have thresholds $t = (n_1 \pm n_4)^2$ which are in our reactions

$$t \rightarrow gN$$
 $t = .4$, .8 Gev^b

The amplitudes of (4.6.3) must satisfy there

$$A = \sqrt{2} \propto t^{\frac{1}{2}} B$$
 (4.6.4)

where

$$\hat{p}_{01} \ll -\sqrt{-t} \ll AB$$
 (4.6.5)

(4.6.4) is interesting as it relates the sign of A and B which on making minimal constancy assumptions and substituting into (4.6.5) gives a definite sign for Pol-

Taking just t = $(n_1 - n_3)^2$ gives as $\sim <0$ in the physical region, $g_{01} > 0$. If you allow A a linear variation in t and apply (4.6.6.4) at both thresholds you find $g_{01} < 0$ but in any cast it is small being proportional to $\sqrt{-t}$ (t- n_a^2) $\sim (-t)^{3/2}$.

Nowever if the π suffers consultacy $\mathfrak{p}_{(1)}$ is proportional to $(t-n_0^2)/\sqrt{ct}$ and for small t is large and predicted from (4.6.4) to be negative whether one includes $t=(n_1\pm n_0)^2$ of course by allowing the π residues extra variation one can obtain either sign in this case as well.

Finally we note that the comparator itself played little role in the above discussion of $g_{i,j}$ to which it does not contribute as it has no helicity O_{i} coupling. Indeed it does not interfere with the w and only above up in $g_{j,j}$, where it adds constructively to the v to nake $g_{j,0}$ smaller in the forward direction, and in $g_{j,j}$, where it adds destructively to the v to ensure $g_{j,j}$ has the necessary vanishing at v = O_{i} . Also note that the comparator has its dominant contribution as spin flip at the NV evertex and so we may ignore its interference with other v = v = nessure, such as the w, whose most important coupling is non-flip.

(11) **™N → pN**

Here we have so much data that I have been forced to choose between those at similar energies. Also to save time on the computer in large fits I have combined successive bins to get ones about .1 dev in width. This reduces X² as it reduce statistical literations but increases the ratio X²/degrees of freedom. Also the width of the y meson produces much greater systematic differences in Total than the statistical accuracy of the experiment would suggest. Thus it proved necessary to allow de/dt scale factors in about one fifth of the socriments.

I tried three types of fit:

(a) First I took all the data with plan > 4 Gev which gave 120 data points. One difficulty is the 8 Gev p on experiment which has in two experiments oftotal varying from .24 to .39. Theoretically we do not expect the total cross section to be greater than .18 and so this data had to be allowed a scale factor which turned out to be rather large. The w dominates this reaction, e.g. at 8 Gev for $\pi^{\pm}p \rightarrow p^{\pm}p$ it is 2/3 of the cross section in 0 > t > - .1 and % that in 0 > t > - 1. So we must decide on what to add to the w and I tried fits with π, A₂; π, ω; π, ω, A₂; π, ω, c; and π, ω, c, A₃ where c stands for conspirator. The first two fits were unsatisfactory giving Y2 > 180. The 3 and 4 pole fits (although of course the conspirator parameters are determined at t = 0 and so do not represent much extra freedom) were all satisfactory with X2 from 140 → 150. There is insufficient data to get conclusive evidence for conspiracy from the density matrix circuits olthough p_{10} is always negative and the data saress well with the compairacy predictions. The nonconspirational fit were also good but did not represent a smaller number of parameters as the "T resident required extra variation. I have always made the compairant chosen consense at $\kappa^2=0$ as as not to produce a particle, and usually given it a slope of .5 although (fix without the A, preferred a smaller one.

One point of interest is the ω parameters. Although this is the dominant $\gamma P = *$ contribution the data is most insensitive to its energy dependence. Blowver, whether it chooses sense or nonenese, its contribution should wantsh when $\kappa_*^{in} = 0$. There is no particular dip in the data until t $^{in} = -4$ when there is some levelling off of defits and $\frac{1}{2} - \frac{1}{4}$ which is suggestive of a ω dip leavened by a conspirator or Λ_2 background. (This dip made fits with only in and in understandard in .) have tried fits without concelling the ω to vanish where $\kappa_*^{in} = 0$ but allowing the residue function a linear variation in t. In many of these fits the residue function chose to vanish at k = -4, and so I have taken $\kappa_0^{in} = -4$ and $\kappa_0^{in} = 1$. (This goes through the ω bole and has an intercept in agreement with the t = 0 WY and K delaids exstating subject), but see the erratum.

One unfortunate feature of all the fits was the small scale factor (.1) preferred by the π .

(b) Encouraged by the above results it was decided to include the lower energy data. But first the three charge states were considered squarately. This produced the expected results, \mathbf{p}^n and \mathbf{p}^n needed smiller parameters but \mathbf{p}^0 needed smcl less $\mathbf{r}^n \leftarrow \mathbf{exchange}$. The small value of \mathbf{g}_{00} in the \mathbf{g}^0 needed much less $\mathbf{r}^n \leftarrow \mathbf{exchange}$. The small value of \mathbf{g}_{00} in the \mathbf{g}^0 data seems to indicate conspiracy but the agreement with the density matrix elements is inferior in all models (sithough there are serious disrepancies between the 2.7 and 2.7 dev data). Unfortunately I have ignored \mathbf{c}^0 contamination which night be expected $\mathbf{c}^{(2)}$ to produce some difference in normalization near to 0. These runs weeded out experiments of inconsistent normalization and equipping these with scale factors we then combined all the data.

(c) On the combined data over all energies we found good agreement with d@dit, Unfortunately there is quantitative disagreement in the experimental density matrix elements at 2.7 and 2.72 dev. In the p^{*}p state the theory agrees best with the 2.72 dev data which shows p₋₁ increasing with that dropping away for t ≈ -.4 in screenent with the ω dip. The 2.7 dev data shows a consistently lower p₁₋₁. However, the p⁶n state was gareed better with the 2.7 dev data.

Fits were better with conspiracy as they gave a lower $\mathfrak Z_{00}$ and a better shape for $\mathfrak Z_{01}$. However as the high energy data was consistent with no conspiracy this could be the effect of non-asymptotic terms which may be expected, from our earlier discussion, to destroy the predicted extra vanishing of $\mathfrak Z_{01}$. Our model did not give an $e^{\mathfrak Z_{-1}}$ term capable of this but the

 A_1 daughter has singular residues and is of type III and can interfere with the $\tau\tau$. Still one must say that the data is consistent at present over all energies with consulracy.

As the theoretical g_{1-1} was too high at 2.7 GeV (although smisfactory at 2.72 GeV) it may be that we are wrong in attributing all the deviations from π exchange to $\tau P = *$ mesons. The inclusion of the h_1 improved χ^2 somewhat and gave a better g_{3-1} the h_1 daughter will also lover g_{3-1} .

In figures 4.3 to 4.5 we give typical fits to $d\pi/dt$ and the density matrix elements at 2.7 and 8 Gev.

(iii) $KN \rightarrow K^*N$

In $W \mapsto \int R^{-1}W$ we have a fit which is very similar to the shorptive moint. Thus we use the same domainent poles and possibly our conspirator represents the cut given by absorptive corrections. However $X^{-1} \times K^{-1}$ is more interesting in that the domains pole is not that used in the absorptive model which has W and ω exchange. In both models W exchange is a much smaller fraction of the total cross section than in

πN → pN but in Regge theory the ω cannot have a large contribution as the experimental de/dt shows no evidence at present for a dip at t - .4. Also SU, would suggest the ω is small but if the ω decides both to violate SU, and not to vanish when " = 0 (this does not contradict analyticity), I have verified it gives as good a fit as the model to be described and it certainly has a more satisfactory energy dependence. Thus one would expect P and P' to be the other important poles. From exact SU, P exchange is forbidden while P' (associating it with fo) has a reasonable coupling which ought to be larger (from the NN vertex) than the P* (associated with the f'(1500)) which has a lover intercent. The depat distribution is rather broad suggesting a P' choosing nonsense but unfortunately the energy variation of total shows a 1/s2 behaviour.

Now let us consider the three charge states separately as in contrast to $\pi N \to \rho N$ they show significant differences, (a) $K^+ p \to K^{++} p$

The experimentalists have done a Regge type fit to their data finding an effective $\mathbf{w}_0 = .26 \pm .27$ but this includes the w-contribution which is, for example, at 2.97 GeV one half the total cross section in 0.9 t.> ... 1 and one quarter that in 0.9 t.> ...

An interesting feature is \mathfrak{z}_{10} which is even positive in some experiments and so in accordance with our previous discussion

fits with a conspiring π were hopeless. Thus I have tried some shaple π + effective τP = + meson fits. On 105 data points we had satisfactory fits for κ_0 = .2 and .4 with χ^2 = 100 and all the disagreement at lower energies. For κ_0 = .69 (P) χ^2 was about 20 higher.

It is perhaps significant that this is the only reaction, out of those I have considered, which prefers moncomspiracy as it is also the one that, due to a lack of direct resonances, has least absorative corrections.

(b) Charge exchange $K^*n \rightarrow K^{0*}p$ $K^*p \rightarrow \overline{K}^{*0}n$

For the exchanges poies I will consider, it requires $g - \Lambda_2$ interference to prosuce a difference between the cross sections for the two charge exchange resctions and this will be a small effect. Thus from isospin conservation the π is such larger than for $K^* P = K^* P$ and the previously containst T = 0 mesons do not contribute. For example, even at 10 GeV the π and combined t P = * contributions to π_{total} are roughly equal.

As in $w^-p \rightarrow g^0n$, g_{00} is rather small and so, unlike x^0 p, compaired is definitely favoured. The g and A_b both give rather small contributions due to their weak NS coupling but the insufficient data and the effect of the TP = 0 compairator obscure any attempt to find a preference for one or the other.

The shape of de/dt at 4.1 and 5.5 Gev is rather broader

than expected theoretically (indeed all three charge states prefer a larger scale factor (.25) for the " π than " π " \rightarrow pN but this may be an effect of the nearer threshold) while the cross section at 10 Gev is about 30% lower than expected.

The variation of the density matrix elements with t is so different at 4.1 and 5.5 GeV that no simple theory could fit them.

(c) $K^-p \rightarrow K^{\bullet-}p$

This charge state contains more data and a larger energy range thin the other two and accordingly 1 have steepide more fits (70) than in any of the other reactions I have considered. $f_{10} \ \ \, \text{is negative for small t and although comparatoris1 fits are matisfactory they prefer a smaller ratio <math>\frac{1}{2} \left(Y_i \right) \left(\frac{1}{k} \right)$

As we sentioned before the energy dependence of \P_{total} does not suggest the high intercept of the \mathbb{P}^* . Finding that \mathbb{P} parameters from a fit to all the data we subtracted off the \mathbb{W} and then fitted the remainder to a fora proportional to \mathbb{P}^* \mathbb{P}^* frective \mathbb{P}^* finding

where we averaged over the given t intervals to reduce the error. On dropping the accurate data in the 2 to 7 Gev range the value of *errective in 0 > t > - .3 was unaltered but the value for the second range changed to - .7 * .12.

We tried fits to Ψ and an effective $\nabla P = *$ meson of even signature with various intercepts and slopes. High values of m_0 (a.4.) were only compatible if they and steep (1.5) slopes but the best fit occured with $m_0 = .2$ and $m_0' = 1.25$ which had a χ^2 of 500 on 180 data points. On adding a compirator with a small alone (c.5) which give a significant contribution for large k, the necessity of associating steep slopes with high intercepts disappeared and the best fit occured with $m_0' = .4$, $m_0' = 1$ and a χ^2 of 400.

On dropping the data between 2 and 3 GeV the assume of W necessary dropped by some 205. (This was like $WN \to gN$ where the lower energy data needed more VP = v than the W provided). Also as the most accurate experiments were now in the middle of the energy range the fits could pivot on these and on varying the intercept between 1 and 7. X^2 only varied by 20 with a minimum of 200 for $K_{\rm cp} = v$. (and 120 data points).

We may conclude that the reaction is probably dominated by a P 'type Regge pole whose parameters are rather sensitive to the method of analysis employed but whose preferred average intercept is nearer .4 than .69.

(d) All Charge States

We then tried a few runs on the combined dats. This was probably not very useful as the separate charge states analyses had suggested very different W parameters for the three cases. Thus K^{*}p did not went conspiracy while K^{*}p although consistent with no compiracy needed a W helicity I coupling of the opposite sign to K^{*}p. Finally K^{*}0 of effinitely wanted conspiracy, we included nothing in our fits to resolve these discrepancies but deylet was fitted reasonably well with the K^{*}p too large theoretically at the higher energies. The residue at the pole agreed well with that expected from the K width. (We had similar agreement for wh * p N although the width seems less well determined exercisements)

on 320 data points our X^2 varied from 1000 for π , P, and Λ_p exchange to 820 for π , c, P, ω and the D meson. We maded the D meson in an attempt to lower the theoretical \mathcal{P}_{1-1} at low energies and also 3 hoped it might reduce the amount of π mecosary leaving, as the D meson is $\pi = 0$, a smaller predicted charge exchange cross section. It succeeded in the first aim but did not have a large enough contribution means to 0 for the second hope to be realised.

In the above fits the contribution of the $\infty \sim$ one fifth that of P^1 and one would expect it to be small as the near equality of the K^{-1} and K^{-1} cross sections suggest little P^1 , ω interference which is of opposite sign in these two reactions.

Due to the above mentioned discrepancies ρ_{10} which usually gave a small contribution to χ^2 now gave as much as the other density matrix elements.

The difference between K and p production is illustrated in figures 4.6 and 4.7.

4.7 Double Resonance Production

the lower threshold 2 constraints

4.7 Double Resona
(i) Constraints

In Section 4.5 on N^* production we only had $\tau P = +$ exchange. We must now consider $\tau P = -$ exchange. This has at

$$J_{5}^{-}(Y_{\frac{1}{2}\frac{1}{2}}^{+}-Y_{-\frac{1}{2}\frac{1}{2}}^{+}) = Y_{\frac{1}{3}/2}^{+}\frac{1}{2}-Y_{-\frac{3}{2}/2}^{+}\frac{1}{2}$$

$$J_{5}^{-}(Y_{\frac{1}{3}/2}^{+}+Y_{-\frac{1}{3}/2}^{+}\nu) = Y_{0}^{+}\nu + Y_{-\frac{1}{3}\nu}^{+}$$

$$(4.7.1)$$

where all amplitudes $\sim 1/[t - (m_2 - m_4)^2]^{\frac{1}{2}}$.

There are at the top thresholds four constraints which we as usual ignore. (The difference in the number of constraints at the two thresholds is because the lighter particle is a fermion according to our earlier discussion about the change in the effective parity for ? states).

Experimentally, $\tau P = -$ mesons appear to have less spin flip than the $\tau P = +$ which may be connected with the reduced number of constraints.

(ii) General Discussion

These reactions are interesting as they provide opportunities to test factorization as the spin structure at each vertex should be the same as in the single resonance production processes.

 $w^*p\to p^{0_k^{n+1}} \ \ \text{has been considered in detail by T. W. Rogers.}$ He may remark here that $\ \ \ y_0 \ \ \text{is negative as one expects at the}$ very least from factorization and our WN + yN fits. There is such difficulty with the normalization and the behaviour of defids near t = 0 that it is difficult to carry the factorization

The only reaction of this kind, that we have made quantitative fits to, is $K^{+}p \to X^{0}N^{+}+^{+}$.

(111) K*p + K^dN*++

. Like the charge exchange reaction in single resonance (K^*) production this isolates the π exchange contribution as being dominant. Some $\tau P = +$ exchange is however necessary and ρ

and \mathbf{A}_2 are available. Both improve the fit obtained with π sione: g by decreasing the theoretical value of f_{00} near t = 0 and A_2 by increasing the cross section for large t.

t = 0 and λ_2 by increasing the cross section for large t. ρ_{10} is a lawys negative not the W parameters necessary for the KK vertex are in rough accordance with those found in K^{*}p and K^{*}on production. The available data does not enable one to distinguish conspiracy as even at 5 fev the physical region boundary is still at t = -.05 and so we have no data at small t. There is no spresens that the positive ρ_{10} of K^{*}p which we tentatively associated with the lack of direct resonances in the K^{*}p system. As there is only one K^{*}p experiment incompatible with negative ρ_{10} perhaps $K^*p + K^*p$ will eventually turn out to agree with the other three reactions supporting negative ρ_{10} .

Compares, from the discussion of Section 4.5, would suggest diffut vanishing at t = 0. There is no evidence in the published data for a turnover in defict and in ny fits without comparing I find good agreement with the expected W residue at its pole. This contrarts with $W^{-1}_{p} \rightarrow \hat{p}^{0.8}_{M^{-1}}$ and indeed whereas the dorn terms for this and $\tilde{X}^{-1}_{p} \rightarrow \tilde{X}^{0.8^{-1}}$ and indeed whereas the dorn terms for this and $\tilde{X}^{-1}_{p} \rightarrow \tilde{X}^{0.8^{-1}}$ and indeed whereas the dorn terms for this and $\tilde{X}^{-1}_{p} \rightarrow \tilde{X}^{0.8^{-1}}$ and indeed whereas the dorn terms for this and $\tilde{X}^{-1}_{p} \rightarrow \tilde{X}^{0.8^{-1}}$ and indeed whereas the dorn terms for this and $\tilde{X}^{-1}_{p} \rightarrow \tilde{X}^{0.8^{-1}}$ and indeed whereas the dorn terms for this and $\tilde{X}^{-1}_{p} \rightarrow \tilde{X}^{0.8^{-1}}$ and indeed whereas the dorn terms for this and $\tilde{X}^{-1}_{p} \rightarrow \tilde{X}^{0.8^{-1}}$ and indeed whereas the dorn terms for this and $\tilde{X}^{-1}_{p} \rightarrow \tilde{X}^{0.8^{-1}}$ and indeed whereas the dorn terms for this and $\tilde{X}^{-1}_{p} \rightarrow \tilde{X}^{0.8^{-1}}$ and indeed whereas the dorn terms for this and $\tilde{X}^{-1}_{p} \rightarrow \tilde{X}^{0.8^{-1}}$ and indeed whereas the dorn terms for this and $\tilde{X}^{-1}_{p} \rightarrow \tilde{X}^{0.8^{-1}}$ and indeed whereas the dorn terms for this and $\tilde{X}^{-1}_{p} \rightarrow \tilde{X}^{0.8^{-1}}$ and indeed whereas the dorn terms for this and $\tilde{X}^{-1}_{p} \rightarrow \tilde{X}^{0.8^{-1}}$ and indeed whereas the dorn terms for the definition of $\tilde{X}^{0.8^{-1}}$ and $\tilde{X}^{0.8^{$

Again one should be able to enforce the $\, g \,$ and $\, \Lambda_2 \,$ couplings from the $\, \, m^{N} \,$ and $\, K^{N} \,$ analyses respectively. However it will be less interesting than for $\, \omega \, ^{N} \, ^{N+*} \,$ as the combined $\, g \,$, $\, \Lambda_2 \,$ contribution is only about 20% that of the $\, m \,$.

Table of Experimental Data Used

Reaction	Reference	Lab. momentum	No. of events	No. of de/dt values	No. of sets of density matrix element values
π ⁺ p →π ⁰ N*++	P.L. 7, 125 (1963)	1.59	115	11	0
	P.L. 13, 341 (1964)	2.79	70	12	0
	P.R. 136, B195 (1964)	3.54	100	20	0
	P.R. 138, B897 (1965)	4	180	14	4
	P.L. <u>19</u> , 608 (1965) and P.L. <u>22</u> ,533 (1966)	8	7	8	1
K ⁺ p → K ⁰ N ⁺⁺⁺	P.R. <u>142</u> , 913 (1966)	1.96	290	7	1
	John Hopkins Univer- sity Preprint	2.26	240	7	2
	N.C. 36, 1101 (1965)	2.97	180	10	3
	N.C. 46, 539 (1966)	3.5	180	- 5	0
	N.C. 46, 539 (1966)	5	. 7	. 4	0
K*p + K ^{O*} N***	P.L. <u>6</u> , 62 (1963)	1.96	280	9	0
	N.C. 39, 417 (1965)	2.97	500	11	5
	N.C. 46, 539 (1966)	3.5	650	5	4
	N.C. 46, 539 (1966)	5	500	4	3

Reaction	Reference	Lab. momentum	No. of events	No. of de/dt values	No. of sets of density matrix element values
K ⁻ p → K ^{*-} p	P.R.L. 16, 485 (1966)	2.1	1500	9	9
	P.R.L. 16, 485 (1966)	2.45	570	3	3
	P.R.L. 16, 485 (1966)	2.64	2150	13	13
	P.L. 12, 352 (1964)	2.97	310	5	3
	P.L. 14, 338 (1965)	3.5	160	20	0
	Argonne Preprint	4.1	1000	18	4
	Argonne Preprint	5.5	600	18	4
	Unpublished	6 .	350	25	0
	Unpublished	10	, 7	15	0
K ⁺ p → K ⁺ +p	P.R. <u>142</u> , 913 (1966)	1.96	170	6	1
	John Hopkins Univer- sity Preprint	2.26	220	8	2
	P.R.L. 15, 737 (1965)	2.3	250	25	2
	N.C. 36, 1101 (1965)	2.97	180	10	3
	N.C. 46, 539 (1966)	3.5	430	5	4
	N.C. 46, 539 (1966)	5	7	5	1
K ⁺ n → K ^{+O} p	P.R.L. <u>15</u> , 737 (1965)	2.3	750	25	3

No. of No. of sets of Reference Rection

			values	eleme value
Argonne Preprint	4.1	135	18	4
Argonne Preprint	5.5	100	18	4
				١.

114

K ⁻ p → K ⁰ n ·	Argonne Preprint	4.1 5.5	135	18 18	4
1	Unpublished	10	?	9	1
π"p → 9 "p	UCRL 16877	2.05	250	20	0
	P.R. <u>152</u> , 1183 (1967)	2.14	620	10	3

Purdue Preprint P.L. 24, 112 (1967)

P.3. 145. 1072 (1956)

P.3. 142, 896 (1966)

P.R. 138, 3897 (1965)

and P.L. 22,533(1966)

P.L. 19, 608 (1965)

Unoublished

UCAL 16877	2.05	250	20	
P.R. <u>152</u> , 1183 (1967)	2.14	620	10	
UCRL 16877	2.36	360	20	
P.R. <u>153</u> , 1423 (1967)	2.7	1000	30	
UCRL 15877 average			_	Ε,

UCRL 16877	2.05	250	20	0
P.R. <u>152</u> , 1183 (1967)	2.14	620	10	3
UCRL 16877	2.36	360	20	0
P.R. <u>153</u> , 1423 (1967)	2.7	1000	30	9
UCRL 15577 average	☆ 2.72	2200	0.	12
from 2.05 → 5.22	- 2.72	****		1
P.R. 145, 1128 (1966)	2.88	720	5	4
UCSL 16877	3.22	630	20	0

4.2 450 30

11

2.08 850 30

4

160

530 2.9

14

10

Reaction	Reference	Lab. momentum		No. of de/dt values	No. of sets of density matrix element values	
π ⁻ ρ → 9 ⁰ n	UCRL 16877	2.05	430	20	0	
	UCRL 16877	2.36	500	20	0	
	P.R. <u>153</u> , 1423 (1967)	2.7	2200	30	6	
	UCRL 16877 average from 2.05 → 3.22	△ 2.72	3500	o	10	
	P.R. 145, 1128 (1966)	2,88	1180	5	0	
	UCRL 16877	3.22	920	20	0	
	N.C. 31, 729 (1964)	4	150	7	0	
	Orsay Preprint	8	160	13	0	

contribution.

 $\mathbf{w}^{\bullet}_{0} \rightarrow \mathbf{w}^{0}_{A}^{0+\nu}$ at 4 GeV. (As in all the figures is/4t is plotted on a logarithmic scale). A fit to \mathbf{p} and \mathbf{p}^{*}_{0} exchange. The solid line represents the \mathbf{p}_{0} contribution and the dashed line the total

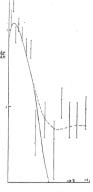


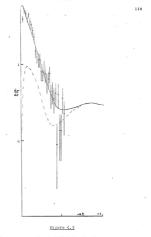
Figure 4.1

rigure 4.3

 $x^{\dagger}p \rightarrow x^{0}x^{0} + \tau$ at 2.97 dev. A fit to A_{2} only. Notice the characteristic difference between this and the p cross section of the previous diagram.



 $\mathbf{w}^{-}>\mathbf{v}_{0}^{-}$) at 2.7 SeV, from a fit to all the prediction data with \mathbf{w}_{1} , \mathbf{w}_{2} and consistency (\mathbf{c}_{1} cackage). The solid line is the solid contribution and the dashed line the \mathbf{u}_{1} and $\mathbf{c}_{1}(\mathbf{v})^{2}+\mathbf{v}_{2}$ contribution. The disappresents on the first point on this such the next diagram is stributed to the former; excellent bits.

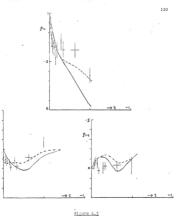


Fi uro 6.4

 $w^{n} \rightarrow \phi^{n}$) at 8 day (notic lines) and $v^{n} \rightarrow \phi^{n}$) at 8 day (dotted lines). That the size fit is the previous diagram with non-invarient ϕ and ϕ giving the same theoretical cross section for these two recetions. Again the solic and desped lines ore the total and $w + \phi$ contributions respectively.



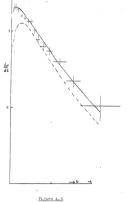
Density matrix elements for $\mathbf{w}^{-} > \mathbf{y}^{-} > \mathbf{y}^{-}$ at 2.7 Over from fits to all the $\mathbf{w} \mathbf{p}$ data. The shift line is fit to \mathbf{w} , \mathbf{w} and \mathbf{h}_2 and the desired line a competencial fit to \mathbf{w} , \mathbf{h}_1 , \mathbf{c} , \mathbf{w} , \mathbf{h}_2 . Notice the characteristic differences in \mathbf{y}^{-}_{10} and the dips in \mathbf{y}_{-1} , due to the \mathbf{w} trajectory passing through sore,



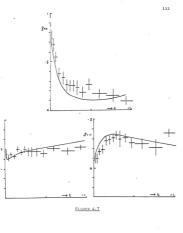
 $K^-p \to K^{0-}p$ at 2.64 GeV from a fit to all the K^0 production data with π , p, c, ω and p^+ exchange. Again the solid line is

the total and the dashed line the c, $\boldsymbol{\omega}$

and P' contribution.



The density partix elements from the same reaction and the same fit as the previous diagram. In g_{10} and g_{1-1} one sees the \mathbf{w} for small t and the D for large that the Confinent $\mathbf{v} \mathbf{p} = -$ contribution.



Two mistakes have been found.

(i) In subsection 4.3 (i) it was not emphasized that the improvement in the position of the cut in s following from using a" not s" only occurse for special mass values. Perhaps s better method is to replace $(e^{-i\pi \omega} + \tau)s^{\infty}$ by $e^{-i\pi \omega}$, (s - symposhala) " + T (uthrashala - u)" but this leads to ghosts for certain integral values of & . (ii) In Section 4.6 the - parameters do not agree with those following from factorization and the elastic scattering data. This analysis of this latter data suggests that all the ω residue functions vanish at t == -.13. Fits to $\pi X \rightarrow \rho X$ which enforce this behaviour agree well with the high energy data with one important exception. This is the occurate 4.2 Gev $\pi^- g \to g^- g$ orth where neither ρ_{00} shows the expected rise nor g_{1-1} the expected dip at t = -.15. However recent $\pi^- p \to p^0 n$ data at 4.2 and 8 GeV confirm the view that the dominant $\tau \dot{r} = +$ contribution to $\pi X \rightarrow \rho X$ is T = 0 exchange. This should be the ω as from SU_{τ} the coupling vanishes.

CHAPTER 5

Electromagnetic Corrections to the Strong Interactions

5.1 Introduction

The tonics to be discussed in this chanter been no relation to the previous work in this thesis. Indeed, for these tonics. we will avoid the complications due to spin, and will consider

There are many excellent accounts in the literature of the theory of combining the strong and electromagnetic interactions. However they use rather different, and in some cases unfashionable, methods. We can senarate four approaches.

(i) Low-energy Here notential H3) and N (N1) methods may be used. The nonrelativistic problem of the NA method has been solved by Corwille and Martin C2). This will provide a good basis for any N/D approach as an N/D calculation is only expected to be reliable at low energies. We will consider this in Section 5.4. which is on Dashen's D2) calculation of the nen mass difference, where we will relate his method to these standard C2) results.

(ii) High-energy Elastic Scattering at t = 0

only spinless non-identical particles.

By this we mean the theory B3),R1),S6) necessary to extract

the "pure nuclear" amplitude from experimental measurements. Ne will consider it in Section 5.5 and translate the previous methods into on-mass shell language. Also we compare it with other situations (e.g. the absorptive model) where one is faced with the problem of combining two potentials.

(iii) General Relativistic Theory

This is expounded in the massive work by Tennie, Frustchi and Suwra⁽¹⁾. Here it is shown bow one can extract the infrared divergent factors from matrix elements in such a way that experimental quantities are finite. They also prove that under certain circumscemes (e.g., large angle scattering but not unfortunately⁽⁵⁾ under the conditions of (iii) this gives the dominant redulative corrections at high energy.

(iv) Rigorous Results

In (iii) the infra-red divergence was treated by using the standard field theory of a massive photon whose mass λ tends to zero at the end of the calculation. In this limit, and working to all orders in e^2 , the amplitude for the production of a finite number of photons is zero. However summing over all possible numbers of photons gives finite experimental quantities which can usually be calculated, as shown in (iii), from the result to first order in e^2 . In a treatment which gives the photon a zero mass from the start, it is unsatisfactory not to have a finite amplitude (as opposed to finite cross sections) and so one can follow $(num_p^{(1)})^{2}$, $(num_p^{(2)})^{2}$, $(num_p^{(2)})^$

and define scattering amplitudes between states containing an infinite number of photoms. Although it is very pleasing to treat the electromagnetic interaction to all orders in e², it seems unlikely that any practical schemes will be based on such an approach. Thus Section 5.2 suggests that the analyticity of the amplitude, to all orders in e², is unsatisfactory whereas, to any finite order in e², it can be treated by the methods usually applied for non-zero mass particles. So it would seem sufficient to treat the amplitude to first order in e², adding in the corrections due to higher order terms, when they are enhanced by the long-range nature of the Coulomb force. (This would be necessary, for instance, at threshold and at t = 0 in elastic accitaring). We shall adopt this pragmatic approach in the following considerations.

(v) Definitions

In order to discuss the best estimate of the "pure nuclear" scattering amplitude we may define some amplitudes.

The nonrelativistic scattering amplitudes is

$$f = \frac{1}{8\pi \sqrt{n}} T ,$$

where T is as defined in the appendix to Chapter 2. We define the Coulomb scattering amplitude $\mathbf{f}_{\mathbf{c}}$, with no strong interactions present, by

$$r_c = \frac{1}{p} \sum_{\epsilon} (2\ell + 1) e^{i\delta_{\ell}^{c}} \sin \delta_{\ell}^{c} P_{\ell}(\cos \theta_{i})$$

11 11 11 11 11 11

with phase-shifts f^C . Similarly introduce the strong interaction amplitude f_g with phase δ^S and the combined amplitude f_{C+g} with phase δ^{C+g} . Define $\bar{\delta}$ by

$$\xi_0^{C+8} = \overline{\xi}_0 + \xi_0^{C}$$
 (5.1.2)

and put $f_{C+B} = f_C + f_2$ where

$$r_2 = \frac{1}{p} \sum_{\ell} (2\ell + 1) e^{2i \frac{C}{\ell}} e^{i \frac{C}{2}} \sin \frac{C}{2} P_{\ell}(\cos \theta_{\ell})$$
 (5.1.5)

 \mathbf{f}_2 might be our first guess at the "nuclear" amplitude but as we will see later a better choice under some circumstances is \mathbf{f}_1 defined by

$$r_1 = \frac{1}{p} \sum_{n} (2 Q + 1) e^{i \vec{k}_{\parallel}} \sin \vec{k}_{\parallel} P_{\parallel}(\cos \theta_{\mu})$$
 (5.1.4)

Nonrelativistically \mathbf{f}_1 has an infra-red convergent perturbation series and it may be interesting to note that it may be found from

i.e. it is the scattering matrix generated by the strong potential where in both the initial and final states one takes, as basis states, Coulomb "in states", rather than the plane wave states used when there is no electromagnetic interaction. Of course we find $f_{\rm c+s}$ by taking out boundary conditions for both potentials in the ket.

Relativistically we have the extra infra-red divergence associated with photon emission, and f_1 still has this divergence. We will also need the relation $R^{(1)}$

$$f_1 = f_2 - 2ip \int f_c^* f_2 - \frac{d\Omega}{4\pi}$$
 (5.1.6)

We will compare the analyticity of f_1 and f_2 in Section 5.2 where we use potential theory as a guide.

In Sections 5.3 and 5.4 we consider the methods of (ii) and (i) respectively.

I would like to thank J.-K. Storrow for discussions on these problems and his work with Landshoff on the more rigorous aspects.

5.2 Nonrelativistic Analyticity

We will consider the analyticity of f_1 and f_2 in nonrelativistic potential theory. This will not be done rigorously but, as we shall find a rather unpleasant behaviour from our own limited treatment, it seems unlikely that either relativity or a fuller nonrelativistic theory will make this any better. The callifities and nonrelativistic seems (see the constitution of the con

similar to first order in e², as long as one takes care of

relativistic crossed diagrams by permuting s, t and u in the argument below. In particular the distinction between s and t disappears relativistically and, for instance, one can obtain

the same result for the nature of the singularity of \(\frac{1}{72} \)

(\(\sum = \text{a} \) photon) from my monrelativistic result ((i) below).

(** = a-photon) from my monrelativistic result ((1) below or from inserting a one photon, nucleon intermediate state in t channel unitarity.

Analyticity in t

Working to first order in the strong interaction coupling, but to all orders in the electromagnetic interaction, we may

calculate^{N1}

which is (for a zero-mass photon)

$$f_1 = f^2N e^{-\pi i \eta} \hat{\Gamma}(1+i \eta) \hat{\Gamma}(1-i \eta) \left[\frac{y - 2ip}{y^2 + 4p^2} \right] \frac{2i\eta}{i \eta + 1}$$

$$= 2F_1 (1+i \eta)^{1-i \eta} + 1 \frac{2p^2}{y^2 + 4p^2} (1 + \cos \theta))$$

$$r_2 = r^2 N e^{-w} \hat{I} - \left[\Gamma(1+i\gamma) \right]^2 - \frac{\left[p - 2i\eta \right]^{2i\gamma}}{p^{2i\gamma} + 2}$$

$$= {}_{2}F_1 \left(1+i\gamma, 1+i\gamma, 1; t/p^2 \right)$$
(5.2.2)

where

$$\gamma = \frac{N e^2}{2p}$$
 (5.2.3)

. The pole at t = μ^2 has been joined by an infinite number of cuts starting there. In f_2 , we see from (5.2.2) that this gives a behaviour $(t-\mu^2)^{-1-2i}$ to the amplitude. We note that:

- (a) The dispersion integral over the discontinuity of such a function does not converge unless provided with predetermined subtractions. This is done automatically by using a photon mass
- $\lambda^2 \rightarrow 0$ but in fact presents no difficulty 87).

logarithms, each time we add a strong line.

- (b) The pole is somewhat obscured by the cuts and relativistically one must be careful about the definition of the electromagnetic mass and coupling constant changes. This is again specified by separating the pole and the cut with a photon mass λ^2 .
- (c) To first order in e^2 the amplitude $f_2 \sim \frac{\rho_{DX}}{x}$ (x = t ρ^2) and the Landau curve of $\frac{1}{1+\epsilon}$ has degenerated into the
- pair of lines $t = \mu^2$ and $p^2 = 0$.

 (d) The result that the singularity on the Landau curve of
- due to _____ is only logarithmically worse than the original singularity of ____, extends to case when we have any number of strong lines. Thus, as in the non-coulomb case, the singularity on the Landau curve improves by a power §, up to

The diagrams, such as , are enhanced to be larger than O(<) by the long range nature of the coulonb force. Although one may not be able to calculate the detailed coulomb corrections one can home to estimate this enhancement.

In the smallyticity approach, this enhancement manifests itself as the rather singular cut we have just discussed. One can therefore decide that the task of evaluating large coulomb corrections reduces to exhibiting functions which do not have such simular cuts.

As from (5.2.1) we find that $f_1 \sim$ original pole term + a term $\ll \ln (t - \gamma^2)$ we have a reason for believing f_1 to be a better estimate than f_2 of the "nuclear" amplitude.

These results have been generalized to the relativistic case by $Storrow^{1/2}$.

- (ii) Analyticity in s
- (a) Threshold Behaviour

Cornille and Martin^{C2)} have shown that

$$\frac{e^{\frac{i}{\hbar}} \frac{\bar{\xi}_{\ell}}{\sin \bar{\xi}_{\ell}}}{c^2 \rho \prod_{\ell}}$$
 (5.2.4)

has a bounded right hand cut determined by a suitable form of

$$C^{2} = 2\pi\eta/(e^{2\pi\gamma} - 1)$$

$$\prod_{\ell} = \prod_{\lambda=1}^{\ell} (p^{2} + \frac{3\epsilon^{2}c^{4}}{4\lambda^{2}})$$
(5.2.5)

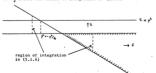
We can now try to investigate the analyticity of the full amplitudes by summing their partial wave series (5.1.3) (5.1.4). We obtain a different result from the non-coulomb case because

the partial wave amplitude no longer has a factor $p^{2\overline{k}}$ concel the $1/2p^2$ in cos $\theta=1$ + $1/2p^2$, δ owe fine that for general fixed it, f_1 and f_2 have an essential singularity as $p^2 \to 0$ while f_1/k^2 has satisfactory analyticity regarded as function of a and cos δ . Never for the particular case to function f_1 and cos δ . Never for the particular case to δ in classic scattering (cos $\theta=1$) we may write a fixed this craim relation.

(b) Difference between f and f

 \mathbf{f}_2 may not be numerically very like the nuclear amplitude (see Section 5.3) but it does at least have expected crossing and analyticity properties.

 \mathbf{f}_1 is defined especially w.r.t. the s-channel, and so cannot have easy s \mapsto u crossing properties. Also it has some unexpected s singularities below threshold which hormally only speer on unphysical sheets. Thus from (5.1,6) we see that \mathbf{f}_1 is not only singular when \mathbf{f}_1 is but also when a t or w singularity of \mathbf{f}_2 enter the domain of integration in (5.1,6).



This is just like a K-matrix or the left-hand cut of a partial wave amplitude and the singularity begins at the same point $p^2 = -\gamma^2/\Lambda$. The discontinuity is trivially evaluated and determines the singularity to be logarithmic. In the language of (i) the cost of removing the singular t discontinuity is this mids a singularity plus a similar remaining logarithmic t singularity. According to the philosophy of (i), both are negligible away from the cut, unless one was using f_1 in a discorreion relation verting to an accuracy of of (s.).

We would like to point out a relation between this unphysical singularity and (5.1.5) which shower f to be secrified by a mixed if prescription i.e. the ket has opposite boundary conditions for the coulomb and strong interactions. Thus in a Lippanna-Rebinger formalism we have to first order-in the alectromagnetic interaction but to all orders in the strong interactions.

$$r_2 - r_8 = 2r_8 \circ r_6 + r_8 \circ r_6 \circ r_8$$
 (5.2.6)

where G is the usual Greens function $\ll \frac{1}{q^2-k^2-i\,\varepsilon}$ while

$$f_1 - f_8 = f_8 (G + G^*) f_C + f_8 G f_C G f_8$$
 (5.2.7)

which has a principal value integral in the first term which

removes the infra-red divergence at the cost of unsatisfactory analyticity.

The crossing and $p^2 = -\gamma^2/4$ difficulty do not occur in f_1 for the leading Regge term in f_- at high energies.

Finally we note that f_1 is also singular when the t singularities of f_0 enter the domain of integration. This indeed is the sook important unexpected singularity and occurs at $p^2 = -\lambda^2/4$ i.e. $p^2 = 0$ in the limit $\lambda^2 = 0$. For the partial wave amplitude it is customary $f^{(2)}$ (cf. (5,2.4/4)) to draw the cut from 0 to ∞ on one opects. One then divides by a factor G^2 in order to ensure that the discontinuity over the new cut is given by unitarity. We will case across this again in Section 5.4.

5.3 Isolation of the Purely Nuclear Amplitude

It is customery to identify f_1 with the "nuclear" mulitude and we must now see under what circumstances this is true. It is certainly not true at low energies, as we see from Section 5.2, and we shall indicate at the end of this section reasons for doubting its usefulness at very high energies.

Using \mathbf{f}_2 has taken account of graphs containing no strong vertices. In order to isolate the amplitude most like that when there is no electrohagnetism (e.g. the amplitude which other by the strong s

which is enhanced above O(≪) by the long range nature of the coulomb force. There is a one to one correspond-

only for a long range force is dominant over a so for a perturbing force of the same

range as the strong force they are of course identical.

One can discuss these problems with any of the classical
opproaches to high energy scattering \$33,113,113 and \$1 x and
Thales \$^{3,1}\$ have noted the relation to the Feynman graph formalism?

 ${
m Thaler}^{{
m Rl}}$) have noted the relation to the Feynman graph formalism ${
m Yl}$. However in order to relate the problem to fashionable schemes, we will use an on-mass shell method.

(i) Nonrelativistic

The difference between f_1 and f_2 has been written out graphically in (5.2.6) and (5.2.7) showing how f_1 still has significant, albeit $O(\ll)$, differences from f_2 .

Work to first order in a and write

when unitarity gives

, Im
$$\delta f = 2p \int \frac{d\Omega}{4\pi} \left(\operatorname{Ref}_{S} f_{C} + \operatorname{Re}(f_{S}^{*} \delta f) \right)$$
 (5.3.1)

- (a) Approximate In \$f by the first term in (5.3.1) and generate se \$f by a fixed t dispersion relation.
- (b) Assume the integrand has no left-hand cut so that we find

' f, 🗠 f ... as desired.

or

Assumption (a) is justified as (5,2)2 makes the second term of (5,5,1) vanish identically. Assumption (b) is the crucial one in that it contains our long range philosophy which has yet to be used. Thus in Section 5.2 we considered the left-hand cut of (5,5,2) showing that it was for the $p^2 = \gamma^2/k$ singularity and not enhanced by the long range force. The most dangerous part of the left-hand cut is that due to the photon pole but it can be treated as mentioned at the end of Section 5.2 by retating the cut to run from 0 to we and using a modified unitarity relation,

(ii) <u>Setimations of f_1</u>

Take the elastic reaction $1+2 \rightarrow 1+2$ at t=0 when the particles have charges z_1e and z_2e .

Then from (5.3.2)

$$f_2 - f_1 = ic \int_{-4\pi^2}^{0} \frac{dt'}{t'-\lambda^2} f_g(s,t')$$
 (5.373)

where $c = \int_{1}^{\infty} z_1 z_2$. (a) Putting $f_S = f_S(s,0) e^{-At'}$ gives R1)

$$f_2 - f_1 = ic log (A \lambda^2) f_g(s,0)$$
 (5.3.4)

which, with $A = (R/2)^2$ and R a nucleon radius, gives Bethe's formula⁸⁵) up to numerical factors inside the logarithm and this is consistent with the neglect of terms which are $O(\ll 1)$.

this is consistent with the neglect of terms which are o(w).

(b) Soloviev^{SO} puts f_s = f_g(s,0) which is sufficient to remove the infra-red divergence but has little else in its favour for this narticular annication. Thereby we get

$$f_2 - f_1 = ic log (\lambda^2/4p^2) f_g(s,0)$$
 (5.3.5)

which has a different energy dependence from (5.3.4). Of course Soloviev's result is quite sufficient if all one wishes is an infra-red finite answer and is not concerned with estimating radiative corrections.

(iii) Relativistic Analysis

The methods used in (i) and (ii) undergo important modifications when we introduce relativity. (a) The change in the photon Born term in (ii) is as usual

(a) The change in the photon Born term in (ii) is as us taken care of by replacing g by its relativistic value

 $\kappa N_{\rm Lio}$. (b) In calculating $\pi \epsilon \delta r$ by a dispersion relation in (i), we must use $f_{\rm S} = \delta r$ to obtain correct relativistic assiyticity. They the extra $f_{\rm S}$ in the mealogue of (5.5.2) has a Left-hand out which must be considered as well as that from the internal contraction.

The fact that

$$P \int_{4n^2}^{\infty} \frac{\sqrt{s'} ds'}{p(s'-s)}$$

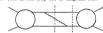
is non-zero whereas

$$p = \left(\frac{ds!}{p(s!-s)} = 0 \right)$$

for s > $4n^2$ is related to the fact that whereas nonrelativistically the infra-red divergent factor is pure imaginary, relativistically we have a real part cancelling with photon emission diagrams in experimental quantities,

(c) Thus to obtain consistent results we must add to (5.5.1) intermediate states containing one photon and any strongly intermetting particles. This is not only due to diagrams such as





which have no infra-red divergence themselves whereas their 2-particle and 2-particle + photon discontinuities are separately divergent.

(5.3.1) becomes

where

$$\begin{split} \mathbf{I}_1 &= & \int_{\mathbf{d}} \mathbf{A}_2 (\operatorname{Re} \ \mathbf{r}_s^{\mathsf{f}} \stackrel{\leftarrow}{\leftarrow} \mathbf{i}_{\mathsf{C}}^{\mathsf{i}} \stackrel{\leftarrow}{\leftarrow} \mathbf{i} + \operatorname{Re} \ \mathbf{r}_s^{\mathsf{f}} \stackrel{\leftarrow}{\leftarrow} \mathbf{i}_{\mathsf{C}}^{\mathsf{f}} \stackrel{\leftarrow}{\leftarrow} \mathbf{r}) \\ \mathbf{I}_2 &= \sum_{n} \int_{\mathbf{d}} \mathbf{d}_{2} \operatorname{Re} \left\{ \mathbf{s}^{\mathsf{f}} \stackrel{\leftarrow}{\leftarrow} \mathbf{n}^{\mathsf{g}} \stackrel{\leftarrow}{\leftarrow} \stackrel{\leftarrow}{\leftarrow} \mathbf{i} + \mathbf{r}_s^{\mathsf{g}} \stackrel{\leftarrow}{\leftarrow} \mathbf{n} \mathbf{s}^{\mathsf{f}} \stackrel{\leftarrow}{\leftarrow} \mathbf{i} \right\} \end{split}$$

and \mathbf{I}_3 is the same as \mathbf{I}_2 except that the intermediate state n has an extra photon. \mathbf{I}_3 can be estimated \mathbf{I}^{11} by making a pole approximation in the variable \mathbf{s}_{12} defined below



to get

$$I_3 \triangleq \sqrt{d \Omega_2 \operatorname{Re}(r_8^{+f} \leftarrow n \quad r_8^{-f} \leftarrow i)}$$

$$\cdot \left\{ \operatorname{g} f \leftarrow i - \operatorname{g} n \leftarrow i - \operatorname{g} f \leftarrow n \right\}$$
(5.3.7)

where the function $\widetilde{\mathfrak{s}}$ is as defined in Yennie, Frautschi and Suura Y1).

If, we try to solve (5.5.6) in the form if * Yf, there is, for our elastic to 0 case, cancellation between \mathbf{I}_2 and \mathbf{I}_3 , and we recover (5.3.4) or (5.3.5) for the coefficient of \mathbf{I}_1 has the unit to know whether to take (5.3.4) or (5.5.5) with their different quantities multisping \mathbf{I}^2 to make it dimensionless. Here there is a real difficulty as this means we must estimate $\widetilde{\mathbf{M}}$ which in reference \mathbf{Y}^{11} would lead to a result like (5.3.5). Although this estimate is presumably wrong 1 do not see how to obtain the same cut-off A in \mathbf{I}_2 and \mathbf{I}_3 .

This may be a real difficulty as the term in $\,$ &f $\,$ & $\,$ &n $\,\lambda^2$

represents a perturbation $d\kappa$ in the Regge pole at $j = \kappa$ we put in (5.3.4). This can be shown as, ignoring signature, a perturbed Regge pole gives an asymptotic behaviour $\delta f \sim \sqrt{4f}$, where

Nowever the rest of $\{5,3,4\}$ with $A \sim \log s$ represents a cut at j = d photon spin - 1 = d with a singular residue. Now it is well known that f cannot have a cut in the f plane and so one may be acceptical about

For the elastic t = 0 case $3^{-1} + 1$ is asymptotically negligible due to s-u cancellation and the trouble only occurs for $n \neq 1$ in (5, 5, 7) and so perhaps (5, 5, 4) holds to higher energies thus in the inelastic case.

The mechanism of the cancellation of the cut in

is probably best examined in a Feynman graph formalism as in Aothe^{R2}).

Finally we note a relation to another theory - the absorptive model - which is meant to correspond to a cancelled cut in the

j plane.
Identify forces as below

(5.3.4).

Absorptive Model		Coulomb	
strong	↔	Coulomb	
w exchange	\leftrightarrow	strong	

as entries on the first line are important only in elastic reactions while those on the second are thus the only inelastic forces. This infancthing of the short and long range forces in unimportant as diagrams such as are symmetric and either force may be regarded as added "externally" to the other.

5.4 N/D Equations

We will now consider how one can include electrosagnetic forces in an N/D equation. We will first treat the problem nonrelativatically pointing out the equivalence between . Dashon and Frautechi's method $^{(2)}$ and that of Gormile and Syrting $^{(2)}$ N. Finally we will consider relativistic effects, For convenience we will only treat a waves and let $a_{\rm g}$ be the ξ = 0-partial wave of the various amplitudes $f_{\rm g}$ introduced in Section 5.1).

(i) Nonrelativistic Treatment

When one considers the S/D equation for the long range coulamb perturbation it is natural, from a naerest singularity philosophy, to suppose that the left-hand cut of a_{c-8} - a_g cun be well approximated by the longest range force, namely single photon exchange, lowever, as we have seen is naillar situations in the previous sections, this is incorrect and leads to infra-red divergences for one must also consider, for instance, whose cut has sowed up to

coincide with that due to As we have said,
using f ₁ estimates to some extent and so
one may distinguish three approaches.
(a) Take $s_0 = s_{c+s} - s_s$ and include as its left-hand cut
plus any diagram used in the original N/D
calculation of a with an extra photon added i.e. if a had
only the Born term as its left-hand cut we would
take + for the left-hand cut of
a _{c+s} - a _s .
The formula of Dashen D1) for the change &s in the boun
state mass ² on adding the coulomb perturbation is found as

Write a = N/D and letting a have a pole R/s - sa we find

$$\frac{1}{2 \text{ w.s.}} \int_{C} \frac{p^{2} \cdot \mathbf{s}_{B} \cdot d\mathbf{s}}{(\mathbf{s} - \mathbf{s}_{B}^{2})} = p^{+2} (\mathbf{s}_{B}) \cdot \mathbf{s}_{B} R \qquad (5.4.1)$$

where c is any counter clockwise contour surrounding s = sa. The integrand has no right-hand cut and so we can convert c into an integral over the left-hand cut. It is easy (see (iii)) to prove (5.4.1) is non-infra-red divergent due to concellation between This is true even when, as above, we calculate a approximately and %a from : } This is because the infra-red divergence of is a function of s with no left-hand cut times

So its left-hand cut involves only In a $_{\rm S}$, and not Re $_{\rm S}$, and it is In $_{\rm S}$ that is calculated exactly in an N/D equation,

The change \$4 in the residue may be considered similarly using the equation

$$\frac{1}{2\pi i} \int_{0}^{\infty} \frac{D^{2} \delta_{B} ds}{(s-s_{B})^{2}} = D^{12}(s_{B}) \delta R + \delta \delta_{B} D^{1}(s_{B}) D^{11}(s_{B}) R \quad (5.4.2)$$

Nowever \S_2 has an instrinsic infra-red divergence which is removed by specifying, for instance, that one should take the residue in f_1 .

In his actual calculation⁰²⁾, as opposed to the theory⁰¹⁾,
Dashen uses a version of this method with, however, a rather
dubious estimate of

- (b) The second way of tackling the problem is to take $\mathbf{a}_1 \mathbf{a}_8$ and only include as the left-hand cut. This is the method advocated by Dashen and Frautschi⁰¹), (c) An entirely equivalent prescription to (b) is that based
- (c) An entirely equivalent prescription to (b) is that based on the rigorous theory of Cormille and Martin⁽²⁾. Namely form $a_{new} = a_1/C^2$ where C^2 is given by (5.2.5) and rotate the one photon left-hand cut to run from 0 to $+\infty$. Then we have

Im
$$\frac{1}{a_{\text{new}}}$$
 = -p C² (5.4.3)

and we get the method (b) on assuming anew has the same left-hand

cut as a .

In order to estimate the error in the approximation (b) or (c) one need only refer to the work of Noyes and Wong^{N1}) where they explicitly evaluate the residual left-hand cut contribution of the land of the contribution of the land o

(ii) Example

In figure 5.1 we plot estimates of \mathbf{f}_{0} for Paton's P^2 J example of the **wave exponential potential. He have taken a potential of strength ** $a/\sqrt{2}$ = 5 (in Paton's P^2 J notation) and give results for N = 1 or 2 where in the unperturbed problem we have taken up to the Nth Born terms. Three different ways have been used for the perturbed problem.

(4) All graphs included which contain one "photon" line and

- up to X strong lines. This is essentially the method (a).

 (a) The neive method using graphs with one "oboton" line and
- u, to N 1 strong lines. This is infra-red divergent. (Y) As (β) but the graph containing N strong lines and an
- (¥) As (β) but the graph containing N strong lines and an external "photon" is estimated as in the method (b) or (c).
 From the figure (X) may appear as good as (S) but this is

probably a spurious effect as, in this model, all the higher order terms omitted in (a) are positive. Thus (X) by overestimating may be accidentally better.

To estimate the error due to a bad unperturbed solution note that with $s_{\hat B} = - q_{\hat B}^2$ we have the following values of



exact		.74
approximate	N = 3	.67
calculation	2	.51
	1	.215

Similar results are obtained from calculations for \$2, However, if one calculates the quantity on the right-hand side of (5.4.2) which is more rapidly convergent than \$s, in (5.4.1), the estimates (€), (\$), (\$) become more nearly equal. But for low N, they are still in disagreement with the exact result showing the rapid convergence of the perturbed problem cannot overcome a bad unperturbed solution.

(iii) Relativistic Treatment

The modifications necessary are very similar to Section 5.7(iii). We will first indicate how one can exhibit the cancellation of the infra-red divergence in \$\$_n and then state which terms cancel among themselves.

. In (5.4.1) we rewrite

$$\int_{c} = \int_{1,h,c.} + \int_{r,h,c.}$$

os on integral over the left- and right-hand cuts. Relativistically is non-zero due to the presence of photon emission diagrams. In the | we replace \$a by \$a' where

\$ n = \$ n' + the perturbed pole terms whose parameters we are trying to find.

We now turn the into an integral over the right-

We now turn the line of into an integral over the right hand cut giving

$$2 \text{ Wi D}^{12}(s_{\underline{0}})$$
 $6 s_{\underline{0}} = \int_{P,h,C_{*}} \frac{ds}{s-s_{\underline{0}}} \left\{ -\ln(6s_{\underline{0}}, 5^{2}) \right\}$ (5.4.4a)
+ Im $(6s_{\underline{0}}, 5^{2})$

This equation is less textological than it movers as in (3.4.4a) δa^{*} is usually evaluated as $\delta a_{1.h.c.}$ (the integral over the left-hand cut of $\delta a_{2.h.c.}$ milet (3.4.4b) is evaluated from unitarity and has contributions only from photon emission diagrams. No have written δa^{*} rather than $\delta a_{1.h.c.}$ as then the infrared divergences will cancel term by term in the integrand. (3.4.4) also allows a trivial extension to the new form of the strip approximation $\delta^{(4)}$ where unitarity is only cufraced for threshold $\delta a_{1.h.c.}$ a strip boundary.

Infra-red divergent terms in (5.4.4) are; in (5.4.4a),

(x) Sc! is the partial wave projection of an infra-red divergent part of . Sf:

- (β) . The one photon Born term has an infra-red divergent partial wave projection;
- while in (5.4.4b),
- (¥) The photon emission contributions to unitarity are infrared divergent,

(a) hes contributions from times a function with are proportional to times a function with a real and imaginary part. The imaginary part, which is all that is present nonrelativistically, cancels with (p). The real part combines with the remaining contributions to (c).



cancel with (%).

We finish with four disconnected remarks on the relativistic

(a) The difference of

$$p\sqrt{s} = \sqrt{(s - (m_N + m_w)^2)(s - (m_N - m_w)^2)}$$

from its threshold value

$$\sqrt{(s - (m_N + m_w)^2) 4 m_N m_w}$$

and in particular the laft-hand cut of p \sqrt{s} need photon emission diagrams to obtain cancellation of the infra-red divergence. This could suggest that photon emission diagrams are important at quite low energies for $\pi N \to \pi N$.

(b) The cancellation of the infra-red divergence will no longer occur for approximations to a_g as happened in the nonrelativistic case. Thus the infra-red divergent part of

Re n_.

(c) Dashen's method calculates \$s_0\$ from an assumption about the strong interactions: namely that the particle is a bound state. The canonical field theory method^{C4} calculates the mass change assuming the special form of the electromagnetic interaction. All graphs of this latter supproach, for examele.



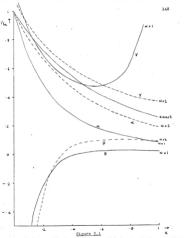


Figure 5.1

 $\mathbf{S}_{8}/\mathbf{S}_{8}$ v. the ratio K of the unperturbed range \mathbf{p} to the perturbed range. \mathbf{S}_{8} is the change in the coupling strength and the other notation is explained in Section 5.4(iii).

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